IX. An account of trigonometrical operations in the years 1821, 1822 and 1823, for determining the difference of longitude between the Royal Observatories of Paris and Greenwich. By Captain Henry Kater, V.P.R.S.

Read January 31, and February 7, 1828.

#### SECTION 1.

In the year 1790, a series of trigonometrical operations was carried on by General Roy, in co-operation with Messrs. De Cassini, Mechain, and Legendre, for the purpose of connecting the meridians of Paris and Greenwich. In England, the work commenced with a base measured on Hounslow Heath, whence triangles were carried through Hanger Hill Tower and Severndroog Castle on Shooter's Hill, to Fairlight Down, Folkstone Turnpike, and Dover Castle on the English coast; which last stations were connected with the church of Nôtre Dame at Calais, and with Blancnez and Montlambert upon the coast of France. An account of these operations will be found in the Philosophical Transactions for 1790.

In the year 1821, the Royal Academy of Sciences and the Board of Longitude at Paris communicated to the Royal Society of London their desire, that the operations for connecting the meridians of Paris and Greenwich should be repeated jointly by both countries, and that commissioners should be nominated by the Royal Academy of Sciences and by the Royal Society of London for that purpose. This proposal having been readily acceded to, Messrs. Arago and Matthee were chosen on the part of the Royal Academy of Sciences, and Lieut.-Colonel (then Captain) Colby and myself were appointed by the Royal Society to co-operate with them.

The instrument employed on this occasion was Ramsden's great theodolite, the property of the Royal Society, the same which had been used by General Roy. A party of the Royal Artillery and a sufficient number of tents were supplied by his Grace the Duke of Wellington, then Master General of the

Ordnance, and every means were afforded which could tend to facilitate the work.

It was at first proposed to adopt some one of the distances given by the Trigonometrical Survey of Great Britain, as a base, and to connect this with General Roy's stations upon Fairlight Down and near Folkstone Turnpike. But the attempt to discover General Roy's stations upon Fairlight and at Folkstone proved, in the first instance, fruitless; and unfortunately, the gun which had marked the important station of the trigonometrical survey on Beachey Head, was not to be found. It is to be feared that, in consequence of some misapprehension, it had been removed along with some old guns which were formerly near that place, and thus one of the most valuable points of the survey of Great Britain was irrecoverably lost.

Colonel Colby was so good as to allow Mr. Gardner (then one of the assistants on the trigonometrical survey, and now agent for the sale of the Ordnance maps,) to accompany us; and to the talents, zeal, and exertion of that gentleman, on various occasions of difficulty, we were much indebted.

The signals used for connecting the stations upon the coasts of England and France were lamps with compound lenses, constructed under the direction of M. Fresnel, and of which he has published an account. It will be sufficient here to mention, that the lens, composed of numerous pieces, was three feet in diameter, and that the light far exceeded that of any of our light-houses, appearing at the distance of forty-eight miles like a star of the first magnitude. Staffs were also erected near the lamps, but these were only occasionally visible.

Having selected convenient stations upon Fairlight Down and near Folkstone Turnpike, and placed the lamps there with steady men to attend them, the party crossed the Channel on the 24th of September 1821, and proceeded to Cape Blancnez, a station to the south-west of Calais. Here we found an old guard-house, the roof of which was partially destroyed, but of which we nevertheless took possession, as it promised a less comfortless abode than our tents at that season would have afforded. At Blancnez we experienced very tempestuous weather; and on the night of the 4th of October it blew so violently that the men's tents were carried away, and we were obliged to take down the theodolite to preserve it from injury.

The observations at Blancnez having been concluded on the 7th of October,

we proceeded to Montlambert (or as it is commonly called Boulambert), a small fort situated on a height near Boulogne; and by the 9th of October the instrument was ready for observing. In the course of our work at this station some delay was experienced in consequence of the lamp at Fairlight not being lighted, and M. Mattheu and Mr. Gardner were dispatched to know the cause of this omission. On their arrival at Calais, finding no packet ready to depart, their anxiety led them to cross in an open boat, at night, in weather so tempestuous that they were nearly lost. They found that the glass chimneys of the lamp at Fairlight were all broken; but their ingenuity remedied this by joining the remaining pieces together; and on the evening of the 13th the light was seen, and satisfactory angles obtained between it and the other stations.

On the 14th of October, the observations at Montlambert being completed, we left that station for Calais. On the 17th we re-crossed the Channel, and on the 19th proceeded to Fairlight. Here I endeavoured to find General Roy's station, and discovered the cause of the failure of the former attempt. In the account of General Roy's operations, his station is stated to be 347 feet southward from the Mill; and the angle at his station between the Mill and Fairlight Church is given. Now it happens that the mill which stood in General Roy's time has been destroyed, and another built upon the Down in a different situation. circular trace however of the old mill was at length discovered; and the distance from its centre to the station having been carefully measured nearly in the proper direction, a small theodolite placed at the end of this radius was shifted until the centre of Roy's Mill and Fairlight Church subtended the given angle. On digging under the theodolite, the wooden pipe by which General Roy had marked his station was found at the depth of four feet. In order to preserve this point, a millstone having the words "Roy's Station" cut upon it, was placed level with the surface of the ground, its centre being precisely over the centre of the pipe.

The observations at Fairlight were completed by the 22nd of October, and the party proceeded on the 24th to a station chosen near Folkstone Turnpike.

In order to carry on the series towards London, stations had been selected on Stede Hill and Wrotham Hill; but as these were not visible from Folkstone, it became necessary to employ an intermediate point on Tolsford Hill.

A staff had been erected on General Roy's station upon Dover Castle, in

order to connect this with the Church of Nôtre Dame at Calais. But as it would have been peculiarly inconvenient, and would have been attended with some risk to have got the great theodolite upon the Castle, the angle there was not taken; but the distance between Dover station and Nôtre Dame has been determined by means of two sides and the included angle, in a manner which will probably appear to be sufficiently satisfactory, as no other station is dependent upon this distance.

The observations at Folkstone were completed on the 27th of October; and with great regret we now bade adieu to our much-esteemed companion M. Arago, who left us for Paris; and as the season was too far advanced to admit of any further proceedings, the party returned to London.

It was now our intention to connect our triangles with the base measured by General Rov upon Hounslow Heath; but though upon examination it was found that the guns marking the termination of this base still existed, it was not thought advisable to attempt to avail ourselves of it, from the many buildings which intervened, and which prevented one end of the base being seen from the other. We were therefore under the necessity of employing the distance from Severndroog Castle to Hanger Hill Tower, as these were the nearest stations to General Rov's base that could be identified with sufficient precision.

During the operations of 1821, I was strongly impressed with the inconvenience of changing the zero point of the theodolite, in order to obtain the angle upon different arcs, so as to do away errors of division; and on my return to London I caused four additional microscopes to be adapted to the instrument, by Mr. Carey. On this important alteration I shall have further to remark in the Appendix.

The summer of 1822 was employed in the choice of stations, one of which was the temporary meridian mark erected near Chingford for the Royal Observatory. This station was chosen, in order that a side of one of our triangles might coincide with the meridian of Greenwich, and that the azimuths of the different stations, with respect to that meridian, might thence be deduced with greater accuracy than might have resulted from observations of the pole star.

Stations were also selected upon Leith Hill, Wrotham Hill, Stede Hill, and Crowborough. We anxiously sought a station to the south of Chingford, for

the purpose of joining it with Severndroog Castle, in order to connect these points with the Royal Observatory; but our endeavours were without success, and we were obliged to content ourselves with accomplishing this object by intersecting the north-west pinnacle of Westminster Abbey, and also the Cross of St. Paul's. The different methods, however, by which the distance from Chingford to Severndroog has been obtained, and the small difference in the results, leave little reason to fear any error of importance.

On the 12th of August 1822, the party proceeded to Hanger Hill Tower. This station is very unfavourable for observations, in consequence of the unsteadiness of the building. Here we felt the great advantage of the additional microscopes with which the theodolite had been furnished, as by their aid we were enabled to accomplish that in a week which it would otherwise have required a much longer period to have completed satisfactorily.

On the 17th of October, our observations being concluded at Hanger Hill, the party left that station for Fairlight Down. Lamps were employed at the stations on Wrotham Hill and Tolsford, and the observations at Fairlight were completed by the 28th.

From Fairlight we proceeded to Folkstone Turnpike. Here, as most of the required angles had been obtained the preceding autumn, little remained to be done, and we were enabled to quit this station on the 5th of September.

Before the party left Folkstone, an attempt was made to discover General Roy's station; and at length the pipe which marked it was found in a state of complete decay, at the distance of three feet to the North-west. The angle between Roy's station and Fairlight being 80° 13′.

The party now proceeded to Tolsford Hill, a commanding eminence, from which the stations on the French coast are visible. Here we saw Fiennes, and succeeded in obtaining the angle between it and Montlambert.

From Tolsford we proceeded on the 9th of September to Stede Hill, a station in the grounds of William Baldwin, Esq. To this gentleman we were indebted for the most kind and friendly attentions, and it would be difficult to do justice to the warm hospitality which we experienced from him. Not only was every thing that could facilitate our objects instantly supplied; but the personal comfort of the whole party, including that of the private soldiers, provided for with the kindest solicitude.

Our observations at Stede Hill being completed on the 14th of September, we left that station on the 16th for Crowborough, and on the 25th proceeded to Leith Hill, a remarkably fine commanding station.

We left Leith Hill on the 5th of October for Wrotham Hill, where, having completed our observations, the party proceeded to Severndroog Castle upon Shooter's Hill. We had here to erect a shed upon the summit of the tower to cover the instrument: this was speedily accomplished by the kind assistance afforded by Lieut. Colonel Jones of the Royal Engineers, who supplied us with carpenters and all that was necessary from Woolwich.

On the 24th of October the theodolite was safely hoisted by proper tackle to the summit of the tower, and the flagstaff having been removed, the instrument was placed with its centre precisely over the spot which the flagstaff had occupied. A platform of boards was attached to the brickwork, so as to be clear of the leads upon which the instrument rested: so unsteady, however, was this building, that we thought it advisable ultimately to reject the angles which had been taken by reading off the five microscopes, in consequence of the disturbance which was found to be occasioned by any person moving upon the platform. We therefore resolved to content ourselves with reading the two opposite microscopes, which might be done without any change of position in the observers. The angles, however, which were deduced from the observations with the five microscopes are given in the Appendix, but are separated by a line from the results furnished by the two microscopes, from which they differ but little, and which have been employed in preference.

At this station we experienced considerable difficulty in obtaining the requisite angles with Hanger Hill, as the signal erected upon that tower was seen only once, in consequence of the intervening smoke of London. At length Colonel Colby thought of a method by which this difficulty was overcome. Tin plates were nailed to the staff upon Hanger Hill Tower, the plates being disposed above each other in certain angles, so as to reflect the sun's rays to Severndroog. This contrivance, which answers the purpose in a certain degree of the heliostat of Professor Gauss, was perfectly successful; each plate gave in succession a neat image of the sun resembling a fixed star, which was seen through a smoke so thick that even the hill was invisible.

From Severndroog the party proceeded to the station at Chingford, and by

the 10th of November the instrument was ready for observation. The season, however, was so far advanced that it was found impossible to obtain the requisite angles with the Royal Observatory or with Westminster Abbey. The health of the men too began to suffer from their being encamped upon a wet clayey soil; we therefore thought it prudent to strike our tents on the 18th, and return to London.

Colonel Colby intending to use Chingford as one of the stations of the Trigonometrical Survey of Great Britain, the theodolite belonging to the Ordnance was placed at Chingford in July 1823, and with it the angles were obtained which we were not able to observe the preceding autumn. This instrument is in every respect similar to that belonging to the Royal Society, excepting that Colonel Colby had recently caused three equidistant microscopes to be adapted to it, which may be used instead of the two microscopes formerly employed.

The transit-room not being visible from Severndroog Castle, the staff erected upon the Royal Observatory was placed upon the centre of the octagonal room of that building; and the angle at Chingford between the staff and the centre of the transit instrument, as well as their distance from each other, is calculated from data furnished by the Astronomer Royal.

As the preservation of the stations was felt to be an object of considerable importance, a stone was procured for each about one foot square and four or five feet long. This was sunk endways until it was level with the ground, and had the word "Station" and the date of the year cut upon it. We did not however rely wholly upon the stone, though its great weight would render its removal a task of some difficulty; each station, wherever practicable, is also fixed by angles formed by steeples or other permanent objects in the vicinity, and by means of which, should the stone be removed, the station may be readily recovered within a very few inches.

## Section 2.—Of the method of computation employed.

A triangle upon the surface of the earth, the sides of which are small in proportion to the radius, may be considered as a spherical triangle, and the sides may be computed by means of spherical trigonometry. Or, the angles formed by the chords may be calculated, and the spherical triangle be thus reduced

to a plane triangle, of which one of the sides and the angles being known, the other sides or chords may be readily determined. This is the method which has hitherto been employed in the English and Indian geodesical operations.

A third method, which is due to Legendre, is as follows: If from each of the observed angles of a small spherical triangle, one third of the spherical excess be deducted, the sines of the angles thus diminished will be proportional to the lengths of the opposite sides, so that the triangle may be resolved as if perfectly rectilineal. This method, which is beautifully simple and accurate, is usually employed on the continent, and is that of which I shall avail myself on the present occasion.

The excess of the sum of the three angles of a spherical triangle above two right angles, termed the spherical excess, is useful to indicate the degree of reliance which may be placed upon the observed angles. I have therefore given it in a separate column, from which the sum of the errors of the observed angles of any one of the triangles may readily be inferred. It is also necessary, when only two angles of a triangle have been observed, that the spherical excess should be known, in order that one third of it may be deducted from each of these angles to prepare them for calculation. The spherical excess of a triangle may be found in seconds, by adding together the logarithm of any two sides, the logarithmic sine of the contained angle, and the constant logarithm 0.3733260.

## Section 3.—Triangles and distances.

The distance given by General Rov from his station upon Severndroog Castle to that upon Hanger Hill Tower, is 84376.68 feet; but the distance from the station of 1822 upon Severndroog Castle to General Rov's station was 10\frac{7}{8} inches; and the angle between General Rov's station and Hanger Hill being about 47° 23′, we have 0.62 of a foot to be added in order to obtain 84377.3 feet, the distance from the station of 1822 to Hanger Hill.

By the comparison of various British standards of linear measure, published in the Phil. Trans. for 1821, it appears that the standard employed by General Roy for the measurement of the base upon Hounslow Heath differed from the Imperial standard yard; and in consequence it becomes necessary to multiply

General Roy's distance by .0000691 to obtain 5.82, the correction to be added to such distance, in order to convert the feet of his survey into Imperial feet \*. Applying this correction, we have 84383.12 for the distance in Imperial feet from Severndroog Castle to Hanger Hill Tower.

	×			
Hanger Hill	from Severnd	roog Cas	tle, 84383.12 fe	et.
	Observed Angles.	Sp. Excess.	Angles for Calculation.	Distances. Feet.
Leith Hill Station  Hanger Hill  Severndroog Castle	35 23 13.87 83 26 23.60 61 10 24.18	••••	35 23 13.32 83 26 23.05 61 10 23.63	127658.21 144760.96
	180 0 1.65	2.53	ļ.	×
Severndroog Ca	stle from Leitl	n Hill St	ation, 144760.9	6 feet.
Wrotham Station Severndroog Castle Leith Hill Station	65 26 47.68 86 25 58.40 28 7 16.42		$\begin{array}{cccccccccccccccccccccccccccccccccccc$	75014.27 158844.37
	180 0 2.50	2.56	and the second s	
Wrotham Stat	ion from Sever	ndroog C	Castle, 75014.27	feet.
Chingford Station	16 35 1.77 149 26 13.36 13 58 44.20		16 35 2.00 149 26 13.58 13 58 44.42	63488.87 133640.58
	179 59 59.33	0.97	-	
Hanger Hil	l from Leith H	(ill Statio	on, 127658.21 fe	et.
Westminster Abbey Hanger Hill Leith Hill Station	84 59 56.81 17 42 36.62		77 17 27.37 84 59 56.41 17 42 36.22	39809.02 130366.27
		1.20		, y

<sup>\*</sup> The sides of the triangles of the Trigonometrical Survey of Great Britain are, I believe, derived from bases measured by General Rox's standard, and they will therefore require the same correction as that employed above, should it be necessary to convert them into Imperial feet.

# Westminster Abbey from Leith Hill Station, 130366.27 feet.

	Observed Angles.	Sp. Excess.	Angles for Calculation.	Distances. Feet.
Severndroog Castle Westminster Abbey Leith Hill Station	62 33 57.67 17 40 36.85		62 33 57.22 99 45 26.38 17 40 36.40	44601.10 144759.97
		1.35		
Westminster Ab	bey from Seve	rndroog (	Castle, 44601.10	) feet.
Chingford Station Severndroog Castle Westminster Abbey	42 52 10.16 61 33 50.95		42 52 9.96 61 33 50.75 75 33 59.29	63488.87 57648.50
		0.59		
Westminster Abbey Leith Hill Station Severndroog Castle	on from Sever 0 , " 17 40 36.85 62 33 57.67	ndroog C	99 45 26.38 17 40 36.40 62 33 57.22	130367.18 44601.41
		1.35		
Westminster Ab	bey from Seve	rndroog (	Castle, 44601.4	l feet.
Chingford Station Westminster Abbey Severndroog Castle	42 52 10.16 61 33 50.95		42 52 9.96 75 33 59.29 61 33 50.75	57648.90 63489.30
,		0.59	-	
Hanger Hill	from Leith H	ill Station	n, 127658.21 fe	et.
St. Paul's	93 13 3.10 19 21 59.50		67 24 58.34 93 13 2.63 19 21 59.03	45848.20 138042.29
		1.41		

St. Paul's	from Leith Hil	l Station	, 138042.29 fee	et.
	Observed Angles.	Sp. Excess.	Angles for Calcu- lation.	Distances. Feet.
Severndroog Castle	72 24 29.57 16 1 15.75		72 24 29.14 91 34 15.54 16 1 15.32	39967.20 144760.30
		1.30		
Severndro	og Castle from	St. Paul'	's, 39967.20 fee	t.
Chingford Station	39 0 35.11 51 43 19.05		39 0 35.95 51 43 18.89 89 16 5.16	63489.66 49844.30
		0.47		
Leith Hill Stat	ion from Sever	ndroog (	Castle, 144760.9	96 feet.
St. Paul's  Leith Hill Station  Severndroog Castle	16 1 15.75 72 24 29.57	•••••	91 34 15.54 16 1 15.32 72 24 29.14	138042.86 39967.36
3		1.30		
Severndroo	og Castle from	St. Paul'	s, 39967.36 feet	t.
Chingford Station	39 0 36.11 51 43 19.05		39 0 35.95 51 43 18.89 89 16 5.16	63489.91 49844.50
-		0.47		

By the preceding triangles we have the following distances from Chingford to Severndroog Castle.

63488.87 63488.87

63489.30

63489.66

63489.91

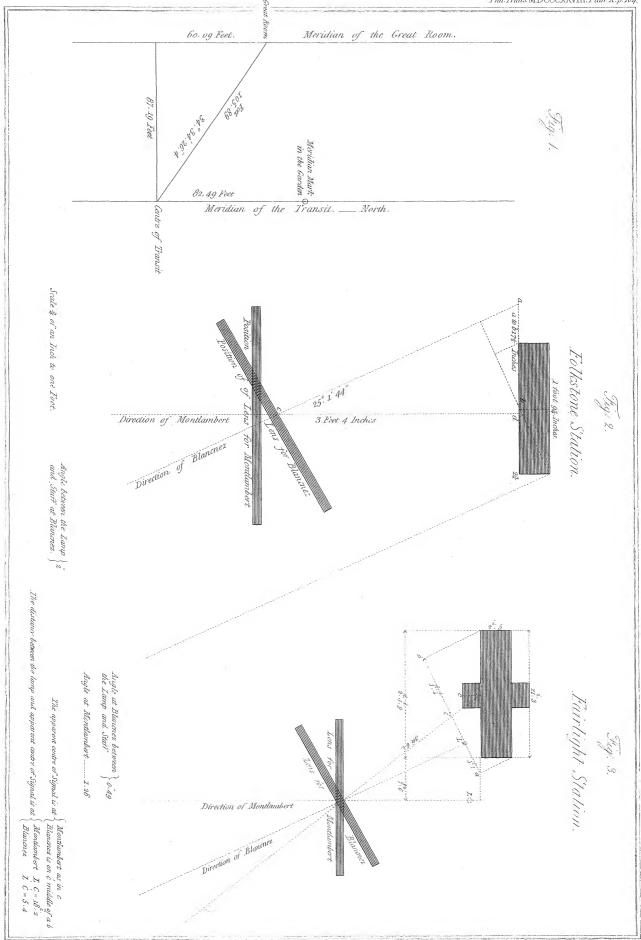
Mean . . . . 63489.32

Mean distance of Severn	ndroog Castle	from Chi	ngford Station,	63489.32 feet.
	Observed Angles.	Sp. Excess.	Angles for Calculation.	Distances. Feet.
Royal Observatory	60 55 21.23 12 51 25.22	0.19	106 13 13.67 60 55 21.17 12 51 25.16	14713.21 57787.63

To connect the centre of the transit instrument at the Royal Observatory with the preceding triangle, the Astronomer Royal favoured me with the data given in Plate X. fig. 1. It may there be seen that the distance from the centre of the octagon room to the centre of the transit is 105.89 feet, the angle at the transit between the octagon room and the meridian of Greenwich 55° 25′ 33″.6, and that the length of a perpendicular let fall from the centre of the octagon room upon the meridian of Greenwich is 87.19 feet. By means of these data and the distance from the centre of the octagon room to Chingford Station, the angle at Chingford Station between the centre of the octagon room and the centre of the transit, is found to be 5′ 11″.21.

If any proof were necessary of the accuracy of the preceding data, I might observe that in the account of General Roy's survey, a plan is given of the Royal Observatory at Greenwich, in which I find the distance from the octagon room to the centre of the transit, and the angle it forms with the meridian, to agree as nearly as possible with the measurements given to me by the Astronomer Royal.

Chingford Station	n from the Roy	yal Obser	rvatory, 57787.6	3 feet.
	Observed Angles.	Sp. Excess.	Angles for Calculation.	Distances. Feet.
Centre of Transit Chingford Station Royal Observatory			55 25 33.6 0 5 11.21 124 29 15.19	57847.66 105.89



With the distance 63489.32 feet, of Chingford Station from Severndroog Castle, the distance 57847.66 feet, from Chingford Station to the centre of the transit, and the contained angle 12° 46′ 13″.95, we obtain the angles and distance given in the following triangle.

=	Observed Angles.	Sp. Excess.	Angles for Calcu- lation.	Distances. Feet.
Severndroog Castle Centre of Transit, Chingford Station			61 3 9.56 106 10 36.48 12 46 13.95	14612.73

The distance given by General Roy from	n t	he	cei	ntre	e of	f tl	1e	
transit to his station on Severndroog,	is	•						14610.58 feet.
Add for difference of stations								0.62
Add to convert into Imperial feet	•	•	•	•	•	•		1.01
General Roy's distance in Imperial feet		•						14612.21
Distance above given	•	•	•	•	•	•		14612.73
	]	Dif	fere	enc	е.			0.52

·	Observed Angles.	Sp. Excess.	Angles for Calcu- lation.	Distances. Feet.
Crowborough Station  Leith Hill Station  Wrotham Station	87 5 15.01 38 56 55.95 53 57 51.13		87 5 14.32 38 56 55.25 53 57 50.43	128615.26 99982.55
:	180 0 2.09	3.03		
Stede Hill Station	on from Crowb 44 44 52.83	orough S	tation, 99982.55	5 feet.
Crowborough Station	41 58 20.94		41 58 19.84	141790.75
Wrotham Station	93 16 49.54		93 16 48.43	94980.95

Stede Hill Statio	on from Crowb	orough S	tation, 141790.7	5 feet.
	Observed Angles.	Sp. Excess.	Angles for Calculation.	Distances. Feet.
Fairlight Station Stede Hill Station Crowborough Station	65 '2 35.83 53 13 24.33 61 44 3.69		65 '2 3'4.54 53 13 23.05 61 44 2.41	137745.52 125267.46
	180 0 3.85	3.70	-	
Stede Hill Sta	tion from Fair	light Sta	tion, 137745.52	feet.
Tolsford Station Fairlight Station Stede Hill Station	69 '7 5'8.69 45 27 54.39 65 24 11.67	•••••	69 '7 57.10 45 27 52.81 65 24 10.09	134038.00 105080.02
	180 0 4.75	3.11		·
Wrotham Static	on from Crowbo	orough S	tation, 99982.55	i feet.
Fairlight Station	33 '6 31.28 43 11 8.83 103 42 24.66		33 6 29.69 43 11 7.27 103 42 23.04	177831.03 125267.87
-	180 0 4.77	2.87		***************************************
Crowborough St	ation from Fai	rlight Sta	ation, 125267.87	7 feet.
Stede Hill Station	53 13 24.33 61 44 3.69 65 2 35.83		53 13 23.05 61 44 2.41 65 2 34.54	141791.20 137745.95
Service	180 0 3.85	3.70		
Stede Hill Star	tion from Fairl	ight Stat	ion, 137745.95	feet.
Tolsford Station	69 7 58.69 65 24 11.67 45 27 54.39		69 '7 57.10 65 24 10.09 45 27 52.81	105080.36 134038.43

	Observed Angles.	Sp. Excess.	Angles for Calcu- lation.	Distances. Feet.
Fairlight Station Wrotham Station Stede Hill Station	31 56 "2.49 50 5 40.62 97 58 17.13		31 56 2.41 50 5 40.54 97 58 17.05	177832.66 137747.27
		3.06		
Stede Hill Sta  Tolsford Station	tion from Fair 69 '7 5'8.69 65 24 11.67		69 '7 57.10 65 24 10.09	feet.
Fairlight Station	45 27 54.39		45 27 52.81	134039.68
	180 0 4.75	3.11		

The preceding triangles give three distances from Tolsford to Fairlight, derived from the three sides of the triangle; Stede Hill, Wrotham, Crowborough,—viz. 134038.00 feet.

134038,43

134039.68

Mean . . . . 134038.70

	Observed Angles.	Sp. Excess.	Angles for Calculation.	Distances. Feet.
Crowborough Station Tolsford Station Fairlight Station	36 5 24.01 33 24 6.89 110 30 29.88		36 5 23.75 33 24 6.63 110 30 29.62	213125.73 125267.72
	180 0 0.78	3.72		
Mean distance Stede	Hill Station fr	om Tolsfo	ord Station, 105	080.58 feet.
		1 1		
Crowborough Station	25 38 39.85		25 38 37.62	
Crowborough Station Stede Hill Station Tolsford Station	25 38 39.85 118 37 36.34 35 43 50.48		25 38 37.62 118 37 34.12 35 43 48.26	141791.16 213127.05

	Observed Angles.	Sp. Excess.	Angles for Calculation.	Distances. Feet.
Folkstone Station Tolsford Station Fairlight Station	36 17 57.06 136 51 46.61 6 50 16.94		36 17 56.85 136 51 46.41 6 50 16.74	26957.63 154811.39
	180 0 0.61	0.58		
Folkstone Stat	tion from Fair	light Stat	ion, 154811.39	feet.
Dungeness Light-House Folkstone Station Fairlight Station	21 14 49.48 21 50 5.80		136 55 "5.46 21 14 49.11 21 50 5.43	84298.42 8213 <b>5.</b> 35
		1.12		,

The following triangles connect our work with the stations on the French coast.

Tolsford Stat	ion from Fairl	ight Stati	ion, 134038.70 i	feet.
	Observed Angles.	Sp. Excess.	Angles for Calcu- lation.	Distances. Feet.
Montlambert Station Tolsford Station Fairlight Station	32 53 2.05 95 48 2.05 51 19 1.68 180 0 5.78	6.07	32 53 0.13 95 48 0.12 51 18 59.75	192717.35 245616.17
Fairlight Statio	n from Montla	ımbert Sı	tation, 245616.1	7 feet.
Blancnez Station Fairlight Station Montlambert Station	75 56 24.49 17 39 26.36 86 24 11.91	• • • • • • • • • • • • • • • • • • • •	75 56 23.57 17 39 25.44 86 24 10.99	252702.93 76800.92
Tolsford Station	180 0 2.76 n from Montla	mbert St	ation, 192717.3	5 feet.
Blancnez Station	103 42 5.81 22 46 47.48 53 31 9.25		103 42 4.96 22 46 46.64 53 31 8.40	159493.81 76803.37
	180 0 2.54	2.81	,	5

Tolsford Stat	ion from Fairl	ight Stati	ion, 134038.70	feet.
	Observed Angles.	Sp. Excess.	Angles for Calculation.	Distances. Feet.
Blancnez Station Tolsford Station Fairlight Station	27 45 39.99 118 34 48.97 33 39 35.99		27 45 38.34 118 34 47.32 33 39 34.34	159500.18 252708.40
	180 0 4.95	4.44		
Tolsford Stati	ion from Blanc	enez Stati	ion, 159500.18	feet.
Montlambert Station Tolsford Station	53 31 "9.25 22 46 47.48 103 42 5.81		53 31 "8.40 22 46 46.64 103 42 4.96	192725.04 76806.43
	180 0 2.54	2.81		
Fairlight	from Blancnez	Station,	252708.40 feet	•
Montlambert Station Fairlight Station	86 24 11.91 17 39 26.36 75 56 24.49	••••	86 24 10.99 17 39 25.44 75 56 23.57	245621.50 76802.60
	180 0 2.76	4.45		
Folkstone Star	tion from Fairl	ight Stat	ion, 154811.39	feet.
Montlambert Station Folkstone Station Fairlight Station	38 44 5'3.42 96 46 26.32 44 28 44.74		38 44 51.92 96 46 24.83 44 28 43.25	173300.64 245618.40
	180 0 4.48	6.29		
Fairlight Station	n from Montla	mbert St	ation, 245618.4	0 feet.
Blancnez Station Fairlight Station Montlambert Station	75 56 24.49 17 39 26.36 86 24 11.91		75 56 23.57 17 39 25.44 86 24 10.99	252705.23 76801.62
	180 0 2.76	4.45		

Folkstone Statio	n from Montla	ımbert S	tation, 173300.6	4 feet.
· · · · · · · · · · · · · · · · · · ·	Observed Angles.	Sp. Excess.	Angles for Calculation.	Distances. Feet.
Blancnez Station Folkstone Station Montlambert Station	107 18 55.90 25 1 46.69 47 39 18.49	2.32	107 18 55.54 25 1 46.33 47 39 18.13	134167.37 76801.57
Folkstone Sta			tion, 154811.39	feet.
Blancnez Station  Folkstone Station  Fairlight Station	31 22 31.13 121 48 13.55 26 49 18.67	4.17	31 22 30.02 121 48 12.43 26 49 17.55	134168.12 252705.93
Folkstone Star	tion from Blan		tion, 134168.12	feet.
Montlambert Station Folkstone Station	47 39 18.49 25 1 46.69 107 18 55.90	2.32	47 39 18.13 25 1 46.33 107 18 55.54	173301.60 76802.00
Fairlight Stat			ion, 252705.93	feet.
Montlambert Station Fairlight Station Blancnez Station	86 24 11.91 17 39 26.36 75 56 24.49		86 24 10.99 17 39 25.44 75 56 23.57	245619.11 76801.82
	180 0 2.76	4.45		2

To show the degree of reliance that may be placed upon the triangles connecting the coasts of England and France, I shall here give the distances resulting from different triangles, derived respectively from the distance Tolsford from Fairlight, and the distance Folkstone from Fairlight.

В	Fairlight.	n By	Folkstone from Fairlight.
Fairlight from Montlambert	Feet. 245616.17		Feet. 245618.40
Turngue nom nomente	245621.50		
Mean	$\frac{245618.88}{}$		<u>245618.75</u>
Fairlight from Blancnez			•
	252708.40		252705.93
Mean	252705.66		252705.58
Tolsford from Montlambert	192717.35		
	192725.04		
Mean	192721.19		
Tolsford from Blancnez	159493.81		
	159500.18		
Mean	<u>159496.99</u>		
Folkstone from Montlambert	 • • • • • • •		173300.64
			173301.60
		Mean	173301.12
Folkstone from Blancnez			134167.37
			134168.12
		Mean	134167.74
Blancnez from Montlambert	76800.92		76801.62
·	76803.37		76801.57
	76806.43		76802.00
	76802.60		76801.82
Mean	n <u>76803.33</u>	• • • •	76801.75

	Observed Angles.	Sp. Excess.	Angles for Calculation.	Distances. Feet.
Dover Station Folkstone Station Blancnez Station	50 37 50.23 12 3 53.88	0.77	117 18 16.41 50 37 49.97 12 3 53.62	31560.06 116726.89
Maan distance Folkst	one Station fro	m Blancr	nez Station, 134	167.74 feet.
Mean distance Polksu		T I		

With the sides "Folkstone to Nôtre Dame, Calais," "Folkstone to Dover," and the included angle 41° 16′ 30″.7, the remaining angles and the distance of Dover Station from Nôtre Dame, Calais, were computed. Also by means of the sides "Blancnez to Dover," "Blancnez to Calais," and the included angle 119° 52′ 50″.32, we obtain another distance from Dover to Nôtre Dame, Calais. These results are contained in the two following triangles:

Folkstone Static	on from Nôtre	Dame, C		
	Observed Angles.	Sp. Excess.	Angles for Calcu- lation.	Distances. Feet.
Nôtre Dame, Calais  Folkstone Station  Dover Station	41 16 30.70		8 42 38.35 41 16 30.44 130 0 51.21	137471.95
		0.78		
Dover Statio	n from Blancı	nez Statio	n, 116726.89 fe	et.
Nôtre Dame, Calais Blancnez Station Dover Station	119 52 50.32		47 24 37.27 119 52 50.04 12 42 32.69	137472.03
	1 3	0.83		

As two of the angles were observed in the triangle "Fiennes, Montlambert, and Blancnez," and as an opportunity occurred at Tolsford of obtaining the angle between Fiennes and Montlambert, I have added the following triangles connecting these stations:

Mean distance	Montlambert	from Bla	ncnez, 76802.54	feet.
î a	Observed Angles.	Sp. Excess.	Angles for Calculation.	Distances. Feet.
Fiennes	34 27 39.83 51 21 31.99		94 10 48.60 34 27 39.62 51 21 31.78	60148.31 43574.27
Mean distance	Montlambert	from Tol	sford 192721.19	feet.
Fiennes	87 58 48.81 17 30 15.75		74 30 57.26 87 58 47.90 17 30 14.84	60148.74 199855.18
	7	2.72		-000000

With the sides "Blancnez from Nôtre Dame, Calais," "Blancnez from Fiennes," and the contained angle, we obtain the distance from Fiennes to Nôtre Dame, Calais.

Fiennes from Blancnez, 43574.27 feet.					
	Observed Angles.	Sp. Excess.	Angles for Calcu- lation.	Distances. Feet.	
Nôtre Dame, Calais Fiennes Blancnez	69 22 53.45	0.34	64 24 14.39 46 12 52.27 69 22 53.34	45221.01	

As we thought it desirable to compare General Roy's operations with our own, staffs were erected upon his stations on Tenterden, Frant, Goudhurst and Lydd steeples; but of these we were able to connect only Frant and Tenterden with our work. The results are as follow:

	Observed Angles.	Sp. Excess.	Angles for Calcu- lation.	Distances. Feet.
Tenterden Church Tolsford Fairlight	0 , " 29 58 10.86 39 19 32.64	•••••	110 42 17.46 29 58 10.38 39 19 32.16	90809.26 71580.75
		1.44		
Fairligh	t from Crowbo	rough, 12	25267.72 feet.	
Frant Church	O , ,,	••••	104 44 17.92	
Fairlight	13 44 20.24 61 31 22.38		13 44 19.97 61 31 22.11	11385 <b>7.</b> 34 30762.92
		0.80		

In a former part of this paper I have mentioned that General Roy's station at Folkstone was discovered at the distance of three feet to the North-west of the new station; the angle between his station at Folkstone and our station at Fairlight being 80° 13′.

At Fairlight, General Roy's station was 87.69 feet to the South-east; the angle between his station and Folkstone being 89° 14′ 31″. The relative positions of the several stations will be better understood from the following diagram,

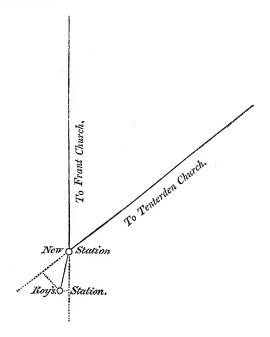


in which R and R designate General Roy's stations, and S and S those of the present operations. From these data the computed distance between General Roy's stations at Fairlight and Folkstone is 154807.00 feet.

We have now several distances which we may compare with those given by General Roy.

The distance from Frant Church to Fairlight is stated by General Roy to be 113928.20 feet. Now if we suppose the distance from Frant to Fairlight to be prolonged, we have the angle between this prolongation and General Roy's station 12° 50′ 56″; and multiplying 87.69 feet, the distance from the new station to that of General Roy, by the cosine of this angle, we obtain 85.48 feet to be subtracted from General Roy's distance, to reduce it to the new station. The distance thus obtained is 113842.72 feet.

In like manner, multiplying 87.69 feet by the cosine of 44° 35′ 41″.75, (the angle between General Roy's station and the prolongation of the distance from Tenterden Church,) we obtain 62.44 feet; which being subtracted from 71634.73 feet, the distance from General Roy's station to Tenterden, will give 71572.29 feet, according to General Roy, for the distance from Tenterden to the new station, without sensible error. The following diagram may serve to render this more intelligible.



General Roy did not obtain directly the distance between his stations at Folkstone and Fairlight; but by using the distances Paddlesworth to Folkstone, Paddlesworth to Fairlight, and the included angle 117° 45′ 42″.65, we are enabled to supply this omission; and we thus obtain 154792.00 feet

for General Roy's distance from Fairlight to Folkstone.—We have also the distance from Dover Castle station to Nôtre Dame, Calais, according to General Roy, 137449.90 feet.

Lastly, The distance from Nôtre Dame, Calais, to Fiennes is given by General Roy, using his own observations, and the angles observed by the French: this distance is stated to be 45219.60 feet.

Converting General Roy's distances into Imperial feet in the manner formerly stated, we have the following results:

From	By General Roy.	By the present Operations.	Difference.
Fairlight to Frant	Feet. 113850.59 71577.24 154802.70 137459.40 45222.72	Feet. 113857.34 71580.75 154807.00 137471.99 45221.01	Feet. 6.75 3.51 4.30 12.59 1.71

Section 4.—Of the distances from the meridian, and from the perpendicular to the meridian, of Greenwich.

It has been mentioned that the station at Chingford was the spot where the temporary meridian mark was erected. This being removed, a staff was put up in its place, having a triangular board fastened to it, the base of which was parallel to the horizon, and the vertex coinciding with the staff.

As it was highly important to ascertain with the greatest precision the situation of this staff with respect to the meridian of Greenwich, Mr. GARDNER went to the Royal Observatory, in order to observe it with the transit instrument. He found that the middle wire of the transit appeared to touch one of the angles at the base of the triangular board, and that the vertex was to the West of the meridian. The angular distance from the meridian to the staff was then measured by means of the micrometer of the transit instrument, and found to be thirty-seven divisions of the micrometer, or 6".16, &c.

By means of the roughly computed distance from the Royal Observatory to Chingford, and its angle with the meridian, the distance of the station from the meridian of Greenwich was found to be 20 inches; and the base of the triangular board proved on measurement to be exactly double that quantity.

When the theodolite was put up at Chingford, the distance of twenty inches was measured to the eastward from the line joining the Station and the Royal Observatory, and an Argand's lamp was placed upon this spot, the position of which I requested the Astronomer Royal to observe. In the Greenwich observations for 1822 I find accordingly, under November 15th, the following remark:—" Observed Captain Kater's light apparently about the thickness of the wire to the west of the meridian." This affords, it is presumed, a sufficient proof that the direction of the station at Chingford, with respect to the meridian of Greenwich, has been accurately determined.

If we suppose a parallel to the meridian of Greenwich and to its perpendicular to be drawn through each station contained in the left-hand column of the following Table, we have the bearings and the distances of the other stations from such parallels, calculated by means of a right-angled plane triangle, the hypothenuse and one of the angles of which are given to find the two other sides: or, let K be the distance between the given stations; M, the distance from the parallel to the perpendicular; P, the distance from the parallel to the meridian; and  $\theta$ , the bearing or angle with the parallel to the meridian. Then,  $M = K \cdot \cos \theta$ , and  $P = K \cdot \sin \theta$ .

TABLE I.

Stations.	Objects.	Bearings.	of Groonwich	
Chingford	Chingford Severndroog Transit Royal Obs. St. Paul's Westminster Abbey Severndroog Wrotham	0 0 6.17 N.W. 73 49 29.69 S.E. 0 0 6.17 S.E. 26 14 15.83 S.W. 30 5 49.84 S.W. 12 46 20.12 S.W. 29 21 22.12 S.E.	1.73 W. 14034.28 E. 1.73 E. 22036.04 W. 28908.97 W. 14035.98 E. 65515.51 E.	57847.66 N. 4070.72 S. 57847.66 S. 44708.80 S. 49876.27 S. 61918.36 S. 116479.69 S.
Severndroog  Leith Hill	Chingford Wrotham Leith Hill Hanger Hill Westminster Abbey St. Paul's Severndroog Wrotham Crowborough Hanger Hill Westminster Abbey St. Paul's	12 46 20.12 N.W. 43 20 6.54 S.E. 43 5 51.03 S.W. 75 43 45.34 N.W. 74 20 11.75 N.W. 64 29 39.83 N.W. 43 5 51.03 N.E. 71 13 6.61 N.E. 69 49 58.14 S.E. 7 42 37.81 N.E. 25 25 14.03 N.E. 27 4 36.84 N.E.	14035.98 W. 51479.66 E. 98906.80 W. 81779.19 W. 42944.95 W. 36072.20 W. 98906.80 E. 150386.41 E. 120729.94 E. 17127.63 E. 55961.13 E. 62835.06 E.	61918.36 N. 54561.77 S. 105703.32 S. 20800.79 N. 12041.68 N. 17209.88 N. 105703.32 N. 51141.52 N. 44341.48 S. 126504.06 N. 117744.81 N. 122912.60 N.

TABLE I. (Continued.)

λ - :		×	÷	1 × 1	10	Distance from	a pa-	Distance from	
=	9.9	- 1			.) .	rallel to the me	ridian	rallel to the pe dicular to the	
Stations.	Objects.		1	Bearings	•	of Greenwi	ch.	dian of Green	
						P. Feet.		M.	. 1011.
						reet.		Feet.	
		0	,	- "		-			
Wrotham	Leith Hill		13		S.W.	150386.41	W,	51141.52	S.
	Severndroog	43	20		N.W.	51479.66	w.	54561.77	N.
* a * * * * * * * * * * * * * * * * * *	Chingford	29		22.12		65515.51	W.	116479.69	N.
	Stede Hill	76	1	32.25	S.E.	92169.87	Ε.	22936.75	$\mathbf{S}_{\bullet}$
	Crowborough	17	15	16.18	S.W.	29656.47	W.	95483.00	S.
Crowborough	Wrotham	17		16.18	N.E.	29656.47	Ε.	95483.00	N.
	Stede Hill	59				121826.34	E.	72546.24	N.
y a	Frant	59	26	16.32	N.E.	26489.28	E.	15642.10	N.
	Tolsford	84		13.64		212272.83	Ε.	19055.17	N.
	Fairlight	59		21.57		107419.64	E.	64443.95	S.
	Leith Hill	69			N.W.	120729.94	W.	44341.48	N.
Stede Hill	Crowborough	59		36.02		121826.34	w.	72546.24	S.
	Wrotham	76		32.25		92169.87	w.	22936.75	Ñ.
	Telsford	59		57.12		90446.52	E.	53491.64	S.
	Fairlight	6		12.97	S.W.	14407.05	w.	136991.06	S.
Fairlight	Stede Hill	6		12.97	Ñ.E.	14407.05	E.	136991.06	N.
Lamingme	Tenterden	12		33.62		15056.77	E.	69979.29	N.
	Tolsford	51	28		N.E.	104853.53	E.	83499.12	N.
*		58		22.52	N.E.	131724.15	E.	81334.62	N.
	Dungeness Lt. House			28.69		1	E.		N.
* .	Blancnez	80				80922.34		14063.10	N.
,	Montlembout	85		40.07		251792.52	E.	21463.18	
	Montlambert	77		54.49	S.E.	239529.39	E.	54353.08	S.
	Crowborough	59	2	21.57		107419.64	W.	64443.95	N.
Talafand	Frant	45		1.60	N.W.	80930.37	W.	80086.04	N.
Tolsford	Fairlight	51		5.78	S.W.	104853.53	W.	83499.12	S.
	Tenterden	81		16.16	S.W.	89797.20	W.	13519.92	S.
,	Crowborough			12.41	S.W.	212272.68	W.	19056.45	S.
	Stede Hill	-		57.12	N.W.	90446.52	W.	53491.64	Ņ.
1.5	Folkstone	85		40.63	S.E.	26870.59	<b>E</b> .	2164.49	S.
	Blancnez	67		41.54	S.E.	146938.81	E.	62034.50	S.
X - 2 14	Fiennes	61		9.18	S.E.	176192.20	E.	94331.42	S.
T 1	Montlambert			54.34	S.E.	134672.78	Ε.	137854.54	S.
Folkstone	Tolsford			40.63	N.W.	26870.59	W.	2164.49	N.
	Dover			20.12	N.E.	28802.86	E.	12900.88	N.
	Nôtre Dame, Calais	-	51	8.14	S.E.	152510.49	E.	47057.49	S.
	Blancnez			49.91	S.E.	120068.38	<b>E.</b>	59871.23	s.
	Montlambert	38		2.31	S.E.	107805.05	Ε.	135688.40	s.
	Dungeness Lt. House	37		33.04	S.W.	50801.68	W.	67271.43	S.
	Fairlight			22.52	S.W.	131724.15	W.	81334.62	S.
Montlambert	Fairlight	77	12	54.49	N.W.	239529.39	W.	54353.08	N.
	Tolsford			54.34	N.W.	134672.78	W.	137854.54	N.
	Folkstone	38	28	2.57	N.W.	107805.22	W.	135688.28	N.
1	Blancnez	9	11	15.56	N.E.	12262.95	<b>E</b> .	75817.22	N.
	Fiennes			55.18		41516.43	Ε.	43522.48	N.
Blancnez	Montlambert	9		16.50	S.W.	12263.29	W.	75817.15	S.
	Fairlight	85		40.07		251792.52	W.	21463.18	S.
	Tolsford	67	6	41.59	N.W.	146938.81	W.	62034.50	N.
	Folkstone	63	29	49.91	N.W.	120068.38	W.	59871.23	N.
	Dover			56.29		91265.50	W.	72772.10	N.
	Nôtre Dame, Calais			50.32		32442.07	E.	12813.75	N.
2	Fiennes			16.34		29253.50	E.	32294.72	S.
	, ,								

From the preceding Table the following is derived, containing the distances from the meridian of Greenwich and from its perpendicular.

TABLE II.

Stations.	Distance from the me ridian of Greenwich. Feet.	Distance from the perpendicular to the mendian of Greenwich.  Feet.
Chinacan	1.73 W.	57847.66 N.
Chingford		4070.72 S.
Severndroog Castle Severndroog Castle	14034.28 E. 14034.25 E.	4070.72 S. 4070.70 S.
		13138.86 N.
St. Paul's	,,	7971.39 N.
Westminster Abbey		58632.03 S.
Wrotham	65513.78 E.	58632.48 S.
Wrotham	65513.92 E.	16730.08 N.
Hanger Hill Tower	67744.93 W.	
Westminster Abbey	28910.69 W.	7970.97 N.
St. Paul's	22037.94 W.	13139.17 N.
Leith Hill	84872.54 W.	109774.03 S.
Severndroog Tower	14034.26 E.	4070.71 S.
Crowborough	35857.40 E.	154115.51 S.
Hanger Hill Tower	67744.91 W.	16730.03 N.
Westminster Abbey	28911.41 W.	7970.78 N.
St. Paul's	22037.48 W.	13138.57 N.
Leith Hill	84872.56 W.	109773.77 S.
Stede Hill	157683.72 E.	81569.00 S.
Crowborough	35857.38 E.	154115.25 S.
Wrotham	65513.86 E.	58632.38 S.
Stede Hill	157683.73 E.	81569.14 S.
Frant Church	62346.67 E.	138473.28 S.
Tolsford	248130.22 E.	135060.21 S.
Fairlight	143277.03 E.	218559.33 S.
Leith Hill	84872.55 W.	109773.90 S.
Crowborough	35857.39 E.	154115.31 S.
Wrotham	65513.86 E.	58632.32 S.
Tolsford	248130.25 E.	135060.71 S.
Fairlight	143276.68 E.	218560.13 S.
Stede Hill	157684.40 E.	81568.67 S.
Tenterden Church	158334.12 E.	148580.44 S.
Tolsford	248130.88 E.	135060.61 S.
Folkstone	275001.50 E.	137225.11 S.
Dungeness Light House	224199.23 E.	204496.55 S.
Blancnez	395069.87 E.	197096.55 S.
Montlambert	382806.74 E.	272912.81 S.
Crowborough	35857.71 E.	154115.78 S.
Frant Church	62346.98 E.	138473.69 S.
Fairlight	143276.92 E.	218559.63 S.
Tenterden Church	158333.25 E.	148580.43 S.
Crowborough	35857.77 E.	154116.96 S.
Stede Hill.	157683.93 E.	81568.87 S.
Folkstone	275001.04 E.	137225.00 S.
Blancnez	395069.26 E.	197095.01 S.
Fiennes	424322.65 E.	229391.93 S.
	-2 10 XX.00	~~J0J1.30 O.

Table II. (Continued.)

	Feet.	dian of Green Feet.	e meri- wich.
Tolsford       248         Dover Castle       303         Nôtre Dame, Calais       427         Blancnez       395         Montlambert       382         Dungeness Light House       224         Fairlight       143         Fairlight       143         Tolsford       248         Folkstone       275         Blancnez       395         Fiennes       424         Montlambert       382         Fairlight       143         Tolsford       248         Folkstone       275         Dover Castle       303         Nôtre Dame, Calais       427	803.23 E. 130.68 E. 804.13 E. 511.76 E. 069.65 E. 806.32 E. 199.06 E. 277.12 E. 276.04 E. 132.65 E. 000.21 E. 068.38 E. 321.86 E. 805.75 E. 276.52 E. 130.48 E. 000.91 E. 803.54 E. 511.11 E. 322.54 E.	272915.05 135060.56 124324.17 184282.54 197096.28 272913.45 204496.69 218559.67 218560.69 135059.23 137225.49 197096.55 229391.29 272913.25 218559.28 135061.60 137224.80 134224.00 184282.35 229390.82	ន់នាន់នាន់នាន់នាន់នាន់នាន់នាន់នាន់នាន់ន

The following Table contains the distance of each Station from the meridian and from the perpendicular to the meridian of Greenwich, obtained by taking the mean of the distances given in the preceding Table.

TABLE III.

	Feet.	dian of Green Feet.	e meri- wich.
St. Paul's       2         Hanger Hill Tower       6         Chingford       6         Centre of Transit, Royal Obs.       -         Severndroog Castle       1         Wrotham       6         Stede Hill       15         Leith Hill       8         Dover Castle       30         Tolsford       24         Folkstone       27         Frant Church       6         Tenterden Church       15         Crowborough       3         Nôtre Dame, Calais       42         Blancnez       39         Dungeness Light House       22         Fairlight       14         Fiennes       42	8910.93 W. 2037.73 W. 7744.92 W. 1.73 W. 4034.26 E. 5513.85 E. 7683.94 E. 4872.55 W. 3803.83 E. 8130.86 E. 5000.91 E. 2346.83 E. 8333.68 E. 5857.53 E. 7511.43 E. 5069.29 E. 4199.14 E. 3276.72 E. 4322.35 E. 805.518 E.	7971.05 13138.87 16730.05 57847.66  4070.71 58632.80 81568.92 109773.90 124324.08 135060.49 137225.12 138473.48 148580.43 154115.76 184282.44 197096.10 204496.62 218559.79 229391.35 272913.64	N.N.N. s.s.s.s.s.s.s.s.s.s.s.s.s.s.s.s.s

Section 5.—Of the latitudes and longitudes of the Stations.

If the earth were a sphere of known diameter, the latitude and longitude of any point upon its surface might readily be calculated by spherical trigonometry. But the earth being an ellipsoid, other methods of computation involving the eccentricity become necessary. This subject has engaged the attention of the most eminent mathematicians, and various formulæ have been given for the purpose of facilitating such computations. These, though equal in accuracy, differ much in practical convenience; and by far the most manageable that I have met with, and of which I shall avail myself on the present occasion, are to be found in a memoir by Oriani, but little known I believe in England, which he published at Milan in 1826, under the title of "Opusculi Astronomici\*."

Let a, be the semi-major axis of the earth, = 3962.439 miles.

b, the semi-minor axis.

e, the eccentricity of the earth = 
$$\sqrt{\frac{a^2-b^2}{a^2}}$$
.

M, the distance in feet from the perpendicular to the meridian at Greenwich.

P, the distance in feet from the meridian of Greenwich.

$$m,=\frac{\mathrm{M}}{b\sin 1''}.$$

$$p, = \frac{P}{b \sin 1''}.$$

L, the latitude of Greenwich.

λ, the latitude of the foot of the perpendicular let fall from the given station on the meridian of Greenwich.

 $\varphi$ , the required latitude of the given station.

u, the required longitude of the given station.

Then I) 
$$\lambda = L \pm m \left[ 1 - e^2 + \frac{3}{2} e^2 \cos^2 \left( L \pm \frac{m}{2} \right) \right]$$

II) 
$$\psi = p (1 - e^2 \sin^2 \lambda)$$

III) 
$$\sin \varphi = \sin \lambda \cos \psi$$

IV) 
$$\tan u = \frac{\tan \psi}{\cos \lambda} \left(1 - \frac{e^2}{2} \cos^2 \lambda\right)$$

<sup>\*</sup> I am indebted for my knowledge of this work to the valuable journal of Baron ZACH.

In computing the eccentricity I have supposed the compression to be  $\frac{1}{300}$ , and I have assumed this (which perhaps for our portion of the meridian may not be very far from the truth), because it is nearly the mean between  $\frac{1}{310}$  and  $\frac{1}{289}$ , the limits between which I believe the ellipticity is generally supposed to be comprised.

Table	of	Latitudes	and	Longitudes.
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Stations.	Latitude.	Longitude.	Longitude in Time.
Chingford Station Hanger Hill Tower St. Paul's (Cross) Westminster Abbey (north- west Pinnacle) Centre of Transit, Royal Ob- servatory Severndroog Castle Wrotham Station Stede Hill Station Leith Hill Station Loiford Station Tolsford Station Frant Church Folkstone Station Tenterden Church Crowborough Station Church of Notre Dame, Calais	51 38 9.59 51 31 22.65 51 30 48.42 51 29 57.34 51 28 38.96 51 27 58.74 51 18 59.35 51 15 7.00 51 10 34.00 51 7 45.59 51 6 8.65 51 5 51.82 51 5 43.18 51 4 5.95 51 3 18.30 50 57 27.95	Congitude.  0 0 0 W. 0 17 51.28 W. 0 5 48.42 W. 0 7 36.95 W.  0 3 41.64 E. 0 17 11.33 E. 0 41 18.86 E. 0 22 12.01 W. 1 19 23.45 E. 1 4 48.19 E. 0 16 16.84 E. 1 11 48.61 E. 0 9 21.45 E. 1 19.20 E. 0 9 21.45 E. 1 51 18.73 E.	
Blancnez Station Dungeness Light House Fairlight Station Fiennes Station Montlambert Station	50 55 29.36 50 54 47.00 50 52 36.88 50 50 4.00 50 43 4.41	1 42 47.45 E. 0 58 18.89 E. 0 37 14.23 E. 1 50 11.41 E. 1 39 9.62 E.	6 51.16 3 53.26 2 28.94 7 20.76 6 36.64

Section 6.—Observations of the pole star for determining the direction of the meridian.

The following is the manner in which observations of the pole star have been usually conducted. The greatest elongation of the star and the time of its greatest elongation being computed, the theodolite was carefully levelled, so that the bubble of the level remained stationary during a whole revolution of the instrument. Then, at the time of the greatest elongation, General Roy states "the angle which the star made with the" (referring) "lamp being noted, the telescope removed, and the plane of the instrument being turned 180° or half

round, the telescope replaced and directed again to the star, the difference on the circle was found to be only  $1\frac{1}{4}$ . The same method was universally adhered to, in all places where observations of the star were obtained."

General Mudge, in his account of the Trigonometrical Survey of Great Britain, says: "At the time of the greatest elongation, when the observer was satisfied of the star being properly bisected, another person at the microscope bisected the dot." "The transit was then taken off, and the instrument being turned half round and the telescope replaced, the star was observed again. This precaution was taken to obviate the errors which might arise from the arms of the instrument being out of the parallel with the plane of the circle, owing to any imperfections in the positions of the Y's on which the transit rested. It was however seldom found that a greater difference subsisted between the readings of the opposite microscopes than what might be supposed to be the consequence of a shake in the centre, or errors in division."

A little consideration will show that the method above described, of obviating an error which might arise from the arms of the telescope not being parallel to the plane of the circle, would not be successful except in the case of the vertical axis being strictly perpendicular to the horizon; but then, the error of the arms of the telescope or axis of the transit, (instantly detected by reversing the level,) could not well escape notice. There is however another source of inaccuracy to which azimuths by the pole star are liable, and which seems to have been wholly disregarded; I allude to an error of the line of collimation. The effect of this upon the azimuth in our latitude would be equal to about six tenths of the error of the line of collimation. This error may however be destroyed by inverting the telescope, or placing that end of the axis which was to the east, to the west; and taking a mean of the observations of the star in both positions.

It must be evident that in taking the greatest elongation of the pole star, the observer is most inconveniently pressed for time; for the azimuth then varies about 1" in four minutes; and besides this, should a passing cloud obscure the star, the observation for that day is lost; consequently by this method of proceeding, a long period is necessary before the direction of the meridian can be obtained.

At Blancnez the weather was so tempestuous that the attempt to deduce

the direction of the meridian in the usual way appeared hopeless, and it occurred to me that it would be a far preferable method to note the time at the moment of observing the star, and thence to calculate the azimuth. I was thus enabled to obtain as many observations as I thought convenient, choosing the time when the star was near its greatest elongation, and when consequently its motion in azimuth was the slowest.

The method pursued was the following:—The instrument being very carefully levelled, some terrestrial object was observed. The telescope was then directed to the pole star, and the star being bisected and the time noted, the microscopes were read off. Several observations of the star having been thus made, the telescope was taken out of the Y's and inverted, the end of the axis which was to the east being now turned towards the west, and sometimes, but not always, the circle was turned 180° or half round. Similar observations of the star were then made with the telescope in this position, and lastly the terrestrial object was again observed.

To lessen the labour of computation, the mean of each two successive observations was taken, and from the calculated azimuth of the star the reading at the meridian was deduced. The mean of such readings for each position of the telescope, compared with the reading of the observed terrestrial object, gave the apparent angle of this object with the meridian; and lastly, the mean of the bearing thus obtained in each position of the telescope gave the true bearing.

On the morning of the 2nd October, 1821, the first observations of the pole star were made at Blancnez; but it blew so violently, that from this, or from some other cause which I cannot discover, these observations, though agreeing well among themselves, differ so widely from those made on the evening of the 3rd, under more favourable circumstances, that I have declined employing them.

The object proposed, in observing the direction of the meridian at Crowborough, Fairlight, Tolsford, and Blancnez, was to obtain the longitude of Blancnez independently of any assumed ellipticity of the earth. Crowborough and Tolsford, and Fairlight and Blancnez are also respectively well situated for obtaining the length of a degree perpendicular to the meridian. It is well known, however, that a very small error in the observed direction of the

meridian will produce an error of considerable magnitude in the length of the perpendicular degree;—and we shall ultimately perceive that deductions of this kind from observations of the pole star, appear to be little, if at all, worthy of confidence.

It has been demonstrated that the sum of the three angles upon a sphere and spheroid is so nearly equal, that the difference when the stations are nearly east and west is absolutely insensible. Having, then, the co-latitudes of two stations, with the observed angle at each, between the meridian and the other station, and consequently the sum of these angles, the difference of longitude or the angle at the pole is obtained by the following method:

As the tangent of half the sum of the co-latitudes is to the tangent of half their difference; so is the tangent of half the sum of the observed angles, to the tangent of half their difference.

The triangle is thus reduced to a spherical triangle, in which two angles and two sides are given to find the third angle.

The deductions from the observations detailed in the Appendix are as follow:

## Crowborough and Fairlight: distance 125267.72 feet.

At Crowborough, the observed angle between the meridian and Fairlight	° 4′	58".36
At Fairlight, the observed angle between the meridian and Crowborough	33	26 .14
The deduced spherical angle at Crowborough 121	13	<b>52</b> .60
The deduced spherical angle at Fairlight 58	24	31.90
The resulting difference of longitude 0	27	46 .67

### Crowborough and Tolsford: distance 213116.39 feet.

At Crowborough, the observed angle between the meridian and Tolsford	84°59′ 34″.35
At Tolsford, the observed angle between the meridian and Crowborough	94 17 21 .56
The deduced spherical angle at Crowborough	84 58 27 .84
The deduced spherical angle at Tolsford	94 18 28 .06
The resulting difference of longitude	0 55 21 .39

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Fairlight and Blancnez: distance 252705.62 feet.

At Fairlight, the observed angle between the meridian and Blancnez	85°	36'	39".73
At Blancnez, (3rd October,) the observed angle between the meridian and Fairlight	93	32	31 .11
The deduced spherical angle at Fairlight	<b>85</b>	<b>35</b>	43 .07
The deduced spherical angle at Blancnez	93	33	27 .77
The resulting difference of longitude	1	5	29

Adding together the longitude of Crowborough and the differences of longitude obtained by means of the azimuths, we have between

Greenwich and Crowborough					•	•					$0_{c}$	9'	21	.45
Crowborough and Fairlight .		•		•				•			0	27	46	.67
Fairlight and Blancnez	,	•	•	•	•	•	•	•	•	•				
Longitude of Blancnez						_								.12

Differing 10".33 in defect, from the longitude found by employing  $\frac{1}{300}$  as the compression.

Section 7.—Of the length of the degree upon a circle perpendicular to the meridian.

I have already remarked that Crowborough and Tolsford, and Fairlight and Blancnez, were respectively very favourably situated for the determination of the length of degrees perpendicular to the meridian at each of these stations: I shall now proceed to state shortly the manner in which the computation was made.

Having obtained by means of the azimuths the difference of longitude, we have a right-angled spherical triangle, the base of which (the co-latitude of the given station,) and the angle at the pole, (the difference of longitude,) are given, to find the perpendicular. Having obtained this arc, we have next to compute the corresponding terrestrial perpendicular. This is effected by means of a small triangle considered as spherical, in which we have the terrestrial distance between the two stations given, and by means of the azimuths two of the angles are deduced. The spherical excess being then computed, and one third subtracted from each of the two angles, the remaining

angle is obtained, and the length of the perpendicular arc in feet is calculated by plane trigonometry.

Lastly, having the perpendicular in arc and also in feet, the number of feet (or fathoms) due to the degree is found by simple proportion.

I shall give the angles of the small triangles used for obtaining the terrestrial perpendicular, in order to facilitate any examination of the work.

A	t Crowbor	At Tolsford.							
Crowborough Tolsford Remaining∠	4 0 25.65 85 42 38.44	<b>ő.</b> 74	85 42	25.40 38.20 56.40	Crowborough	84 59 34.35	<b>ő.</b> 80	84 59 90 43	21.3 34.1 4.6
				•					
	At Fairlight. At Blancnez.								
Fairlight Blancnez Remaining∠	4 23 20.27 86 27 28.89	<b>1.</b> 15	86 27	19.89 28.51 11.60	Fairlight	3 32 31.11 85 36 39.73	<b>ő.</b> 93	3 32 85 36 90 50	39.42

In the manner before explained, we obtain

The perpendicular arc at Crowborough 34' 47".84 equal to 212532.00 feet. The perpendicular arc at Tolsford . . . 34 45 .71 equal to 212329.66 feet. The perpendicular arc at Fairlight . . . 41 19 .32 equal to 252250.29 feet. The perpendicular arc at Blancnez . . 41 16 .77 equal to 251991.98 feet.

The length of the degree perpendicular to the meridian at Crowborough	61077	fathoms.
at Tolsford	61081.3	fathoms.
at Fairlight	61045	fathoms.
at Blancnez	61045.3	fathoms.

And taking the means of the latitudes of Crowborough and Tolsford, and of Fairlight and Blancnez, and the means of the respective perpendicular degrees, we have

The perpendicular degree in lat.  $51^{\circ}$  4' 43''.47 = 61079.17 fathoms. and in lat. 50 54 3 .12 = 61045.15 fathoms.

A moment's examination is sufficient to show that these results are totally unworthy of credit; and that the length of the perpendicular degree above given, must be erroneous about one hundred fathoms.

As very great care was bestowed in making the observations, it is important to determine the degree of error in the azimuth, which would produce an error so considerable, as that which is here indicated.

If 2" be added to the azimuth at Crowborough and to that at Tolsford, the resulting difference of longitude would be diminished 5".14, and the length of the perpendicular degree would be increased 95 fathoms.

Now an error of two seconds in azimuth may proceed from such a variety of sources, that it is scarcely possible to detect it. I think no one acquainted with the great theodolite would venture to assert that the level and its adjustment comprising that of the Y's, can be depended upon to within two seconds of the truth, and an error of 2" in the level would affect the azimuth to the amount of about 2".3. This error arising from the level, I have before explained is not to be destroyed by turning the instrument half round; and were there no other source of inaccuracy, I should consider this alone, as an insurmountable objection to the determination of azimuths by means of observations of the pole star.

But in addition to this, the level may be affected by irregular local density. At Arbury Hill (one of the stations of the Trigonometrical Survey of Great Britain), it is known that the plumb-line of the zenith sector was deflected, so as to occasion an anomaly of  $5\frac{1}{2}$  seconds in latitude. The same cause of disturbance would equally affect the level, and this admits of no remedy.

To the sources of inaccuracy, before enumerated, may be added a small uncertainty, (the fraction of a second for example,) in the polar distance of the pole star, which would influence the azimuth nearly double that quantity. The possibility, and I might perhaps venture to say the probability, of horizontal refraction, affecting the situation of the terrestrial object to which the star is referred, may also be considered; but this last is common to every method of obtaining the direction of the meridian.

From what has been advanced, it should seem that observations of the pole star, for the purpose of determining the length of the perpendicular degree in

our latitude, are wholly unworthy of credit, and that some other method less liable to error should be employed.

Of these, the best which has occurred to me, is the well known method of observing a star when near the east or west point of the horizon, and from the time and the calculated azimuth to deduce the place of the meridian. Here the alteration of the azimuth from a variation in the refraction, must be carefully taken into account, and the altitude of the star must therefore be obtained.

It will not, however, be necessary to observe the star when very near the horizon, as the error in the azimuth arising from the level decreases as the tangent of the altitude, and at an elevation of 12° is scarcely more than two tenths of the error in the horizontality of the axis of the telescope.

Section 8.—Of the heights of the stations above the level of the sea, and of the terrestrial refraction.

Let the arc between the two stations be A. The depressions reciprocally observed at the two stations reduced to the height of the axis of the theodolite be D and d; and let R be the mean terrestrial refraction.

Then  $R = \frac{A - (D + d)}{2}$  and should one of the stations appear elevated from

the other station, calling the elevation E, we have  $R = \frac{(A+E)-D}{2}$ .

The axis of the instrument was about  $5\frac{1}{2}$  feet above the ground, and the angle subtended by this, at the distance between the stations being computed and subtracted from the observed depression, the apparent depression of a point at the height of the axis was obtained. The distance between the stations was converted into arc, by allowing 101.7 feet for each second, and with the arc and the apparent depression, the refraction was computed. The refraction being added to the depression, the difference between this and half the contained arc, gave the angle subtended by the difference in the height of the two stations above the level of the sea; the height of that station being in excess, at which the true depression exceeded one half of the contained arc. Lastly, the angle thus obtained, and the distance between the stations, gave the difference of their heights in feet.

At Folkstone Pier, there is a flag-staff, the height of which was carefully measured, and found to be 37 feet above the Pier. From the pier to the mark on the tide gauge indicating XXI feet, was 5 feet. The harbour master informed me that the highest spring tides rose 20 feet. We have therefore 43 feet from the top of the flag-staff to high water mark, and 63 feet to low water mark.

By means of the side Tolsford from Folkstone in the following triangle, the distances from Folkstone to the flag staff, and from Tolsford to the flag staff were obtained.

Folkstone Pier Flag-staff						•	
Folkstone station			84	37	49	6152.0 fee	t.
Tolsford station			13	4	14	27083.6 fee	t.

With the depression of the summit of the flag-staff, observed at Folkstone and the above data, we obtain the height of the axis of the instrument at Folkstone station above low water mark 559.1 feet.

Similar observations were made at Tolsford, and the resulting height of that station above low water mark differed only 3 feet in defect from that obtained by means of Folkstone. The refraction employed in these computations was  $\frac{1}{13.2}$  of the contained arc.

In the following Table, are given in one view the data and the computed results.

	Observed Depression or Elevation.	Contained Arc A.	Angle subtended by 5½ feet.	tion	$\frac{1}{\frac{A}{R}}$	Angle of Difference of Height.	Differ- ence of Height. Feet.	low
Folkstone Horizon of sea	8 23 37	, ,, 25 33.1	1 11	, ,, 1 56.1		0 1 11		576.9
Folkstone Fairlight	0 10 44.6 0 11 13	25 21.85	0 7.3	1 49.42	1 14	+0 0 14.2	10.5	559.1 569.6
Folkstone Dover Castle	0 13 50	5 10	0 35.6	0 23.5		+0 11 2.9	101.4	559.1 457.7
Folkstone,	0 12 30	21 59	0 8.4	1 39.9		+0 3 2.0	118.4	559.1 440.7
Folkstone	4 40 5	1 0.38	3 4.4	0 4.6,		+4 36 35	496.1	559.1 63.0

	Observed Depression or Elevation.	Contained Arc A.	Angle subtended by $5\frac{1}{2}$ feet.	Mean Refrac- tion R.	$\frac{1}{\overline{\mathbf{A}}}$	Angle of Difference of Height.	Difference of Height. Feet.	Above low Water. Feet.
Folkstone E Tolsford D		4 25	0 42.0	0 23.9	-   1   1 1	+0 3 14.4	25.4	559.1 584.5
Tolsford	0 23 39.7	25 36.7	ų.	1 57.0	-	- -		581.8
Tolsford	$\begin{array}{cccc} 0 & 6 & 10 \\ 0 & 8 & 12 \end{array}$	17 13	0 10.8	1 36.3	$\frac{1}{11}$	+0 1 01.0	31.1	584.5 615.6
Tolsford	1 8 22	4 26.25	0 42.0	0 20.2		+1 5 48	518.5	581.5 63.0
Tolsford	$\begin{bmatrix} 0 & 9 & 5 \\ 0 & 9 & 7.5 \end{bmatrix}$	21 58	0 8.5	2 1.25	1 11	+0 0 1.25	0.65	584.5 585.2
Fairlight	0 23 48.5	25 45.6	, 0	1 57.1		4.		586.7
Fairlight	0 2 22 0 15 29	20 30	0 9.0	1 28.5	1 14	+0 6 33.5	239.0	577.4 816.4
FairlightStede Hill	0 8 43.5 0 10 34	22 34	0 8.2	1 46.45	1 2	+0 0 55.25	36.9	577.4 614.3
Fairlight	0 8 56.5 0 16 35	29 8	0 6.3	1 54.55	1 15	+0 3 49.25	197.7	577.4 775.1
Stede Hill Horizon of sea,	0 23 56	25 53.6		1 57.6	7.			591.6
Stede Hill Wrotham	$\begin{array}{cccc}0&1&0\\0&12&57\end{array}$	15 34	0 11.9	1 0.4	1 15	+0 5 58.5	165.1	615.0 780.1
Stede Hill	$\begin{matrix}0&5&24\\0&14&49\end{matrix}$	23 14	0 8.0	1 38.5	1 14	+0 4 42.5	194.2	615.0 809.2
Crowborough Horizon of sea	0 27 58	30 15.4		2 17.4				813.0
Crowborough Leith Hill	0 4 47.5 0 13 21	21 4	0 8.8	1 37	1 13	+0 4 16.3	159.6	812.8 972.4
Crowborough Wrotham	0 8 9 0 6 19.5	16 23	0 11.4	1 8.6	1 14	+0 0 54.7	26.7	812.8 786.1
Wrotham Leith Hill	0 7 15.5 0 15 17.25	26 1	0 7.1	1 51.2	114	+0 4 0.8	185.6	780.4 966.0
Leith Hill	0 30 42	33 12.9		2 30.9				978.3
Leith Hill	0 21 47	23 43	0 7.8	1 48	-	+0 11 35.7	488.4	966.0 477.6
Blancnez	0 20 33.3	22 14.4		1 41.1				438.2
					.			

At every station from which the sea was visible, the depression of the horizon

was carefully observed, and the resulting heights will be found in the preceding table. These serve to verify to a certain degree the conclusions otherwise obtained.

The mean of the proportion of the refraction to the contained arc, is  $\frac{1}{13.2}$ , and this has been employed on every occasion where the refraction was not deduced from reciprocal observations.

I shall now give the elevation of the ground at each station, above the level of the sea at low water, the point chosen by other observers.

	Above low water. Feet.	By depression of horizon.	Difference. Feet.
Folkstone Station	553.6	571.4	+17.8
Tolsford Station	579.0	576.3	_ 2.7
Blancnez	435.2	432.7	<b>— 2.5</b>
Fairlight	572.0	581.2	+ 9.2
Dover Castle Battle-	452.2	465.8	*****
Stede Hill Station	609.5	586.1	-23.4
Crowborough Station	807.3	807.5	+0.2
Leith Hill Station	960.5	972.8	+12.3
Wrotham Station	775.0		
Severndroog Castle Battlements }	472.1	Mean	+ 1.5

The mean of the differences is so small that we are authorized to conclude that no error of consequence exists in the heights above given; but as the depressions of the horizon were taken probably in various states of the tide, the mean result should perhaps have differed in defect about 10 feet. This, however, may be fairly attributed to uncertainty in the refraction employed.

In the course of the operations which have been detailed, great pains were taken to identify the stations, by the bearings of such objects as were conveniently situated for the purpose;—these are given in the Appendix.

It is to be regretted that our excellent associate M. Arago has not yet published the results of his operations in France; and I must therefore, in the absence of higher authority, take the longitude of Calais, as given in the Connaissance des Tems, to be 0° 28′ 59″ west of Paris. Adding this to 1° 51′ 18″.73 the east longitude of Calais from Greenwich, given by the present work, we obtain 2° 30′ 17″.73 for the difference of longitude between Paris

and Greenwich. This converted into time is 9<sup>m</sup> 21<sup>s</sup>.18, differing only 0<sup>s</sup>.28 in defect from the admirable results obtained by the operations with fire signals, reported in the Philosophical Transactions for 1826, by Mr. Herschel.

The truth of the preceding work wholly depends upon the degree of reliance that may be placed upon the base on Hounslow Heath; and as the accuracy of this is in some measure questionable\*, it is certainly desirable that a new base should be measured, to connect in the most unexceptionable manner the stations at Leith Hill and Wrotham. The measurement of a base has hitherto not kept pace with the progress of other parts of geodetical operations; but the elegant arrangement which Lieut.-Colonel Colby has recently imagined for compensating expansion, and which has already been tried in Ireland with perfect success, leaves no doubt of the future accuracy of this most important part of trigonometrical operations.

#### APPENDIX.

I HAVE reserved for an Appendix such remarks as could not have been introduced in the body of the work, without interrupting the regular connection of its parts.

The original observations are deposited with the Royal Society, and may be consulted whenever occasion may require. It has not been thought necessary to print them, as all the angles employed in this work have been carefully deduced from them, and are given at the end of the present communication. The letter prefixed to each angle indicates the name of the observer; and where the degrees and minutes are repeated, it is to be understood that the instrument has been shifted, and the readings for the angle taken upon different parts of the circle.

The great theodolite had originally only two opposite microscopes, and until the addition I am about to describe was made, the observations were conducted in the following manner.

The instrument being carefully levelled, the objects were intersected, and the microscopes were read off; but it is evident the truth of the angle thus obtained would depend upon the accuracy of the divisions of the circle from

<sup>\*</sup> See Phil. Trans. for 1821, "On the Comparison of various British Standards," &c.

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which it was deduced. In order to do away any error of this kind, the whole instrument was shifted by turning it horizontally a few degrees; and being again levelled, the observations were repeated, and the angle was obtained on different parts of the circle. This operation was repeated seldom fewer than eight times, which it was supposed would be sufficient to do away errors of division. Now as at each observation the angle is deduced from readings taken on four different parts of the circle, eight repetitions of this kind would give a mean angle deduced from thirty-two different points of the instrument. The time, however, required for this was so considerable as to constitute a very serious objection; in addition to which, when the instrument had been recently shifted, it was feared the spring of the parts might introduce error. These inconveniences led me to have four additional microscopes fixed to the theodolite, at such distances as with one of the original microscopes to divide the circle into five equal parts. This arrangement of any number of microscopes or verniers which form a prime number, and the manner of using them, is due to Mr. Pond the Astronomer Royal, but was never published by him. By means of five microscopes, raising the telescope from the Y's, turning the circle 180° in azimuth, and repeating the observations, the angle is obtained upon twenty different parts of the circle, without shifting the instrument, and consequently any error of division may be supposed to be reduced to a very small quantity. Employing in like manner three equidistant microscopes, the angle is obtained by readings upon twelve different parts of the circle.

The second original microscope was not removed, and this afforded an opportunity of comparing the angles obtained by two opposite microscopes, with those deduced by means of five.

In the course of this work I remarked a curious fact, new to me, and for which I was at a loss to account. In hazy weather when the staff was so faint as to be only just visible, it disappeared upon bringing it to the intersection of the cross wires, so that the angle could not be observed.

A remedy for this inconvenience was suggested and put in practice by Mr. Gardner. The horizontal spider's web of the micrometer being moved above the centre, Mr. Gardner succeeded in lodging upon it a very minute particle of dust. When the image of the staff was brought to this, it appeared as if planted upon a mole-hill, and we were thus enabled to observe with great

accuracy. I consider this as a very important improvement in the theodolite, and we availed ourselves of it upon all occasions excepting in the observations of the pole star.

As I was desirous of knowing the degree of precision with which an object could be observed with the telescope of the great theodolite magnifying about fifty times, and also the accuracy with which the microscopes could be read off, as well as the comparative merits of cross wires and Mr. Gardner's dot, we resolved to make some experiments on the subject. A staff upon a steeple was taken which was faintly seen; Colonel Colby marked the time occupied by the observations, and Mr. Gardner read off a certain microscope. The position of the telescope and of the micrometer of the microscope were of course altered between each observation. The following were the results:—

With the Cross Wires.

Time at the commencement ... 0<sup>h</sup> 43<sup>m</sup>.

Observations.	Readings.	Observation	1S.	Readings.
1	$35''\frac{3}{4}$	6.		$36''_{\frac{1}{2}}$
2	. $35\frac{1}{2}$	7.		$35\frac{3}{4}$
3	$35\frac{3}{4}$	8.		. $36\frac{2}{3}$
4	$35^{\frac{2}{3}}$	9.		$36\frac{3}{4}$
5	$35 \frac{4}{5}$	10.		$36\frac{3}{4}$
	Mean	. 36".0	3	

Time at the end . . .  $0^h$   $45^m$   $15^s$ .

With Mr. GARDNER'S Dot.

Time at the commencement ...  $0^h 47^m$ .

Observations.	Readings.	Observation	Readin	Readings.			
1	$1''\frac{1}{2}$	6.		$1''\frac{7}{8}$			
2	$1 - \frac{3}{4}$	7.		. 2			
3	$1\frac{3}{4}$	8.		. 1 7/8			
4	. 2	9.		. 2			
, <b>5</b>	. 2	. 10.		. $2\frac{1}{4}$			
	Mean .	1".9	)				
Time at	the end	Oh	48	m 27s			

The time occupied in making ten observations with the cross wires was  $2^m$   $15^s$ , and the greatest difference from the mean 0''.72.

The time required for ten observations with Mr. Gardner's dot was 1<sup>m</sup> 27<sup>s</sup>, and the greatest difference from the mean only 0".4.

These experiments appear to be important; they seem to show that in any single observation the combined errors of the telescope and microscopes cannot exceed, when the cross wires are employed, three quarters of a second, and when Mr. Gardner's dot is used, they amount only to four-tenths of a second. The time, too, required for the latter observations is little more than half of that which is requisite for the former.

Much error has been supposed to arise both in astronomical and geodetical observations, from unequal expansion of the limb of the instrument. In order

to bring this to the test of experiment, the index of one of the microscopes was placed at zero, and a certain division on the circle brought to its wire. The other microscopes were then read off, the divisions under them having been carefully bisected, and the mean was registered. A piece of lead was placed in boiling water until it acquired the same temperature; and it was then laid upon the limb of the instrument, between two of the microscopes. Having allowed some time to elapse, the first division was again brought to the zero microscope, and the other divisions bisected by their respective micrometers; when the readings were found to be very different, but the mean varied little from the mean first taken. These experiments were repeated with the same results, and satisfactorily proved that no error of consequence is to be feared from unequal expansion of the circle when several microscopes are employed.

Very different, however, was the consequence of applying the hand to any one of the radii to which the microscopes were attached. Then, the expansion which took place immediately and to a very considerable degree, affected the mean of the readings, by altering the position of the microscope to the support of which the hand had been applied.

From these experiments we may infer the very great importance of securing the permanent respective positions of the microscopes. Perhaps this might be best effected by imitating the principle of the mural circle. In this instrument the microscopes are firmly attached to a wall, and any sensible change in their relative positions can scarcely be imagined to take place. In like manner the microscopes of portable instruments might be fixed to a solid plate of metal; and this being a good conductor of heat, should any partial change of temperature take place it is probably to be expected that it would be so rapidly diffused throughout the whole mass as to occasion no perceptible change in the relative distances of the microscopes from each other \*.

During our stay at Fairlight, a source of error was remarked which it may

<sup>\*</sup> In instruments constructed in the usual manner, where the microscopes are attached to arms or radii, these may be covered to some thickness by strips of flannel or leather, and thus the ill consequences to be apprehended from currents of air of different temperatures may, perhaps, be avoided. The great theodolite was treated in this manner; but as this was done at Chingford, the last station we visited, no opportunity was afforded of remarking the effect.

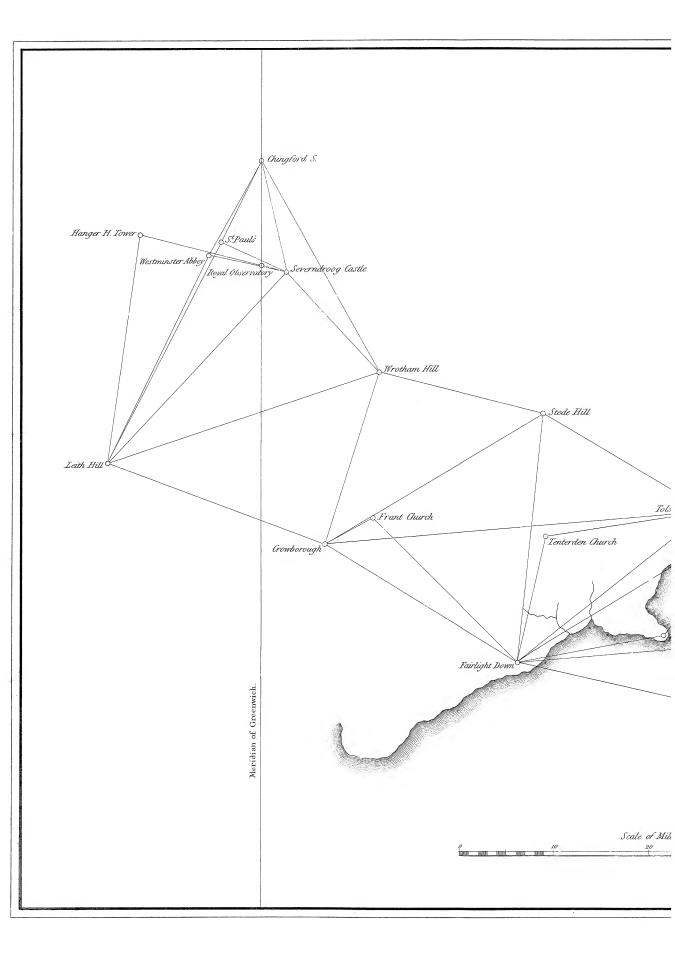
be important to mention. An object having been carefully bisected by moving the tangent screw slowly in one direction, and the microscopes read off, the result was found to differ three seconds from that obtained when the tangent screw was moved slowly in the opposite direction. This could arise only from friction upon the axis, and the yielding of the radii of the circle, when drawn by the tangent screw clamped to its circumference. Numerous experiments were made with similar results; so that the force thus applied to the circumference of the circle, occasioned an error of one second and a half, plus or minus, according to the direction in which the tangent screw was made to act.

On shaking the screw, if I may so express myself, backwards and forwards with little jerks, before the object was finally bisected, the error just described was obviated. It would perhaps, however, be preferable, instead of giving motion to the instrument by means of a tangent screw acting on the circumference of the circle, to have a bar, connected at one end with a tangent screw, and a collar at the other end passing round the axis, to which it might be clamped at pleasure. The axis would then be the first part moved, and the probable error arising from dragging the instrument round by the limb would be avoided. This arrangement seems to be particularly called for in circles of large dimensions.

The errors which may arise from lateral refraction have often been suspected but never clearly ascertained. In the course of our work, however, we had such evidence of the fact as to leave no doubt of its existence. The angle between the same objects would differ under the most favourable circumstances about five seconds on different days, and perhaps a second and a half or two seconds may be considered as the error which may affect an angle from lateral refraction in an ordinary state of the atmosphere. During the observations at Stede Hill one fine day, the telescope being directed to the staff at Wrotham, a shower rapidly approached from the left, and the staff gradually receded from the cross wires until it was obscured by the intervening haze. At Leith Hill, after unfavourable weather, it cleared up in the evening, and though there was no wind, very extraordinary differences were perceived in the angles, for which it would have been difficult to assign any other cause than lateral refraction, varying considerably at short intervals.

At Montlambert we had very decided evidence of lateral refraction affecting the angle between Fairlight lamp and Tolsford lamp. These objects were taken on the evening of the 13th October, by M. Arago, Colonel Colby, and myself, under the most favourable circumstances. The observations were repeated on the morning of the 14th, and the mean of the five deduced angles differed from the mean of the seven angles of the preceding evening 2".13. In the table of the angles at Montlambert I have separated those of the evening of the 13th from those obtained on the morning of the 14th by a dotted line. So persuaded was I of the existence of this source of inaccuracy, that I seldom left any station until I found the difference between an angle observed at different times amount to about five seconds, which I considered to be the extreme limits of error.

There is a source of error in the great theodolite, and which may attach, though from different causes, to other instruments which are read off by means of microscopes; I allude to the variation in "the run", as it is termed, of the microscope. The head of the micrometer is divided into sixty parts, consequently each part should be equal to one second, and one revolution equal to a minute. In order to effect this, the object-glass of the microscope is placed by the observer at such a distance from the limb of the instrument as to form an image of the arc comprised between two neighbouring divisions, of such an extent as that the micrometer head shall make an even number of revolutions equal to the number of minutes, in passing over the space from division to division. This having been nearly attained, the tube containing the eye-glass and micrometer is moved until the divisions are seen distinctly, and the operation is repeated until the result is satisfactory. Now, in taking the instrument from station to station, the cone of the theodolite is lowered upon the axis to prevent injury, and the adjustments just described must be repeated at each station. But this is not the only inconvenience; for the microscopes will frequently alter their run from expansion, or any cause which may affect the distance between the limb and the object-glass. On the occasion of comparing certain standards of linear measure, I abandoned the usual form of the micrometer microscope, for an arrangement which, I conceive, possesses decided advantages over it; and which, in the case of circular instruments, appears to be liable to no other inconvenience than that of re-



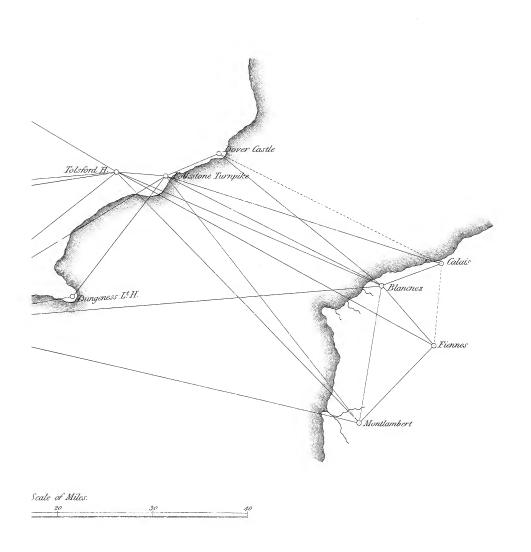
TRIANGLES

for connecting the

MERIDIANS OF PARIS

and —

GREENWICH.



quiring the use of a table to reduce the measurements to arc. If a microscope furnished with cross wires fixed in the focus of the eye-glass be moved parallel to itself by a micrometer screw, it will measure the actual distance which it passes over, and no error can arise from a variation of the distance between the limb of the instrument and the object-glass of the microscope. It possesses also this further advantage, that the object-glass may be changed, and the power of the microscope varied at pleasure without affecting the scale. This construction, I cannot but feel, would be a considerable improvement if applied to astronomical circles, where very minute quantities are the objects of research; as any error arising from the want of strict perpendicularity of the plane of the circle to the axis of motion, or from any other cause which might vary the distance of the circle from the microscope, would be avoided.

I have stated that staffs were erected at Fairlight and at Folkstone near the lamps. These were occasionally taken at Blancnez and Montlambert, and a correction therefore became necessary to reduce such angles to what they would have been had the lamps been observed. To obtain the data for computing this correction, Mr. Gardner made a plan (Plate X.) of the relative positions of the staff and lamp for each station.

A sketch of the triangles constituting the present work is given in Plate XI. The Tables require little explanation. The angles are given as deduced by means of five microscopes, and also by the two opposite microscopes. At Blancnez the letters indicating the names of the observers have not been prefixed, as the angles were individually determined by every one of the party. To the observed angles at each station is added the manner in which such angles have been derived, as could not be obtained by direct observation. Lastly, the readings are given of such objects as were selected for the purpose of identifying the stations: and here it is necessary to bear in mind that the degrees of the great theodolite are numbered from zero to 180°, and that then the numbering is recommenced.

Tables are also given detailing the observations of the Pole-star. From the observed time, the error of the chronometer, and the time of the star's southing, the horary angle is computed in mean time and in degrees; and by means of the horary angle, the polar distance, and the co-latitude, the azimuth is deduced, and the reading at the meridian is obtained.

### Tables of Observed Angles, &c.

## At Hanger Hill.

S	everndroog aı	nd Lei	th Hill.		Westminster Abbey and Leith Hill.					
Observers.	Five Microscopes.	Mean. Five.			Observers.	Five Microscopes.	Mean. Five.			
G. G. G. K.	83 26 23.85 21.25 83 26 24.90 24.30 83 26 23.90 23.50	23.70 23.92			G. G. G. K. K.	84 59 59.16 57.75 84 59 56.95 57.65 84 59 59.90 56.75 84 59 59.87 57.95 84 59 58.95 84 59 54.80 56.55 54.40 57.35 54.83	58.45 57.30 58.32 58.91 58.95 55.67 55.87			
	Mean			+	G.	52.15 54.37				
Reje	cting the last,	23.60				Mean	56.81			
]	Leith Hill and	St. I	Paul's.						1	
Observers.	Five Microscopes.	Mean. Five.								
G.	93 13 2.50 3.70	" 3.10			-					

### Observations for identifying the Station.

	Readings.		ngs.	. **		eadi	ngs.
	0	- 1	11		0	1	11
Harrow Spire	178	48	35.31	Dome of Chelsea College	136	13	51.50
Small white Spire, distant about				The Ball of the Horizontal Mill			
two miles, supposed Kingsbury	<b>52</b>	12	37.50	at Battersea	144	35	14.50
Hendon Church Vane-staff	<b>57</b>	5	49.12	Battersea Spire	145	17	20.50
Pancras New Church	111	56	11.75	The Centre of Knockholt Beeches	154	17	12.00
Marylebonne Church	114	20	<b>51.5</b> 0	Cupola of a new Church, five miles	158	33	59.25
Bayswater Chapel	116	8	11.25	Croydon Church Tower	165	20	51.62
St. Paul's	118	20	56.90	Kew Pagoda	21	46	56.00
Westminster Abbey (N.W. pinnacle)	126	34	7.21	Vane of Ealing Cupola	36	31	56.25
Leith Hill Station	31	34	6.12	Windsor Castle Flag-staff	102	33	39.75

At Severndroog Tower.

7	Wrotham,	and Leit	h Hill.	*		Chingford and	d Wro	otham.	
Observers.	Five Microscope	Mean. Five.	Two Mi- croscopes.	Mean. Two.	Observers.	Five Microscopes.	Mean. Five.	Two Microscopes.	Mean. Two.
K. C.K.& G.	57. 58.	.70 " .10 57.20 .30 57.77	58.12 57.00 58.82 58.87 59.38	" 57.91 59.12	K. C.K.&G.	149 26 18.75 17.75 18.10 16.10 15.60	" 18.25 17.10	17.88 17.50 18.63 14.12 14.13	17.69 16.37
G. K.	56 57 52	.45 .96 .55	63.25 58.92 56.00 57.44	57.46	G.	18.50 21.20 18.60 16.37	17.05 19.90	18.62 15.38 24.75 15.00	16.37 20.06
<b>C</b> .	58 57 56	.30   56.20 .00   56.75	60.00 58.88 57.00	58.72 57.94	K.	18.10 15.75 149 26 13.25	16.92 12.60	14.87 15.25 12.63	15.06 11.81
G. K.C.&G.	62 86 25 56		59.62 62.25 52.50	60.93 54.25	,	11.95 13.50 12.25	12.87	11.00 13.50 13.75	13.62
	54 50	.70 .90 .65 52.77	56.00 54.12 48.88	51.50	· .	149 26 14.45 18.45	16.45	14.12	15.00
		.65 .10 54.37	52.87 53.25	53.06	C. & K.	149 26		11.79 10.13 10.00	10.96
K.C.&G.	86 25 86 25	7	57.00 59.31 53.75	58.15 56.87	*	149 26	•	12.00 17.00 20.00	11.00
			60.00 57.75 59.38	58.56		149 26		12.00 11.62 14.37	11.81 13.56
	86 25		57.12 58.38 59.50	57.75 59.25		149 26		12.75 15.75 13.88	14.81
	86 25		59.00 60.00 60.37	60.18	1.	149 26		12.75 14.00 17.13	13.37 14.63
		50 56.65	59.50 56.00 57.00	57.75 57.87		149 26 15.15 15.05	15.10	12.13 13.38 15.13	14.25
140°	54 58	.80 .95 .90 56.92	58.75 57.50 58.25	57.87	G. G.	16.10 15.50 15.45	15.80 15.17	16.00 15.25 10.37	15.62 12.62
	57	.65 01.02	59.88 60.38 56.75	60.13 58.06		149 26		14.87 13.75 15.13	14.44
General	58 mean	.40	59.38	57.67	K. & G.	12.85 11.50	12.17	16.50 10.00	13.25
	low the line			58.40		elow the line			14.31 13.36

<sup>\*</sup> A bad dot.

# Severndroog Tower (Continued).

I.	Ianger Hill a	nd Ch	ingford.		Chingford, and Westminster Abbey.					
Observers.	Five Microscopes.	Mean. Five.	Two Mi- croscopes.	Mean. Two.	Observers.	Five Microscopes.	Mean. Five.	Two Microscopes.	Mean. Two.	
к.	62 57 24.05 23.30	" 23.67	2 <sup>"</sup> 5.25 23.62	" 24.44	К.	61 33 50.25 52.65	" 51.45	50.13 52.87	″ 51.50	
K.C.&G.	62 57		24.50 25.75	25.12	C.	48.55 50.90	49.72	48.37 50.88	49.62	
			22.50 21.38	21.94	K.	51.95 49.55	50.75	53.13 49.25	51.19	
	62 57		22.75 23.25 21.87	23.00	G.	52.00 51.80	51.90	52.50 52.50	52.50	
- "	02 31		23.87 24.50	22.87	-	Mean	50.95	_	51.20	
	62 57		23.12 26.50	23.81 26.00	Chingfo	ord, and Gree	nwich	Observa	tory.	
*	÷		25.50 25.27 27.19	26.23	Observers.	Five Microscopes.	Mean. Five.	Two Microscopes.	Mean. Two.	
G. G.	62 57		23.38 24.63 25.00	24.81	C.	60 55 19.10 20.75	″ 19.92	1 <sup>"</sup> 9.37 22.50	″ 20.93	
7.	Mean		23.67	24.25		19.30 24.05	21.67	~~.00		
	Leith Hill an	d Chi	ngford.	•	K.C.&G.	20.05 20.50 23.25	21.87	21.00 22.50 25.75	24.12	
Observers.	Five Microscopes.	Mean. Five.	Two Microscopes.	Mean. Two.		20.60 22.35	21.47	23.50 23.12	23.31	
к.	124 7 42.55	"	44.00	"		Mean	21.23		22.79	
C.	44.80 46.60	45.70	44.37 47.06	45.71		Chingford an	d St.	Paul's.		
C. C.	46.25 44.10 42.35	45.17	47.00 42.00 41.37	44.50	Observers.	Five Microscopes.	Mean. Five.	Two Microscopes.	Mean. Two.	
G. K.	45.67 47.80 45.95	46.87	46.08 47.69 44.75	46.22	G.	51 43 17.20 19.35	" 18.27	1.6.87 19.38	" 18.12	
K.C.&G.	51.35	51.02	54.87 53.00	53.93	K.	19.33 18.15 18.85	18.50	18.88 19.50	19.19	
K.C.&G.	51.60 57.10	54.35	42.38 57.37	49.87	C.	15.65 21.20	18.42	16.00 21.75	18.87	
	Mean 47.43	48.62	47.07	48.05	K.	20.55 19.95	20.25	21.63 21.62	21.62	
		Ē	1		G.	20.25 19.40	19.82	21.12 21.00	21.06	
						Mean	19.05		19.77	

### Angles derived at Severndroog Castle.

Chingford and Wrotham	$1\overset{\circ}{4}9$	26	13.36
Wrotham and Leith Hill	86	25	58.40
Hanger Hill and Chingford	62	57	24.25
Sum	298	49	36.01
	360	0	0.00
Hanger Hill and Leith Hill	61	10	23.99
Hanger Hill and Chingford	62	57	24.25
Chingford and Leith Hill	124	7	48.62
*	61	10	24.37
	61	10	23.99
MeanHanger Hill and Leith Hill	61	10	24.18

Chingford and Leith Hill Chingford and Westminster Abbey			
Leith Hill and Westminster Abbey	62	33	57.67
Chingford and Leith Hill			
Chingford and St. Paul's	51	43	19.05
Leith Hill and St. Paul's	72	24	29.57

## Observations for identifying the Station.

	R	eadi	ngs.	× .	F	Readi	ngs.
Eltham Spire	$\overset{\circ}{145}$	27	51.25	West Cupola of Greenwich Hos-	0	1	H
Cupola on Bromley Tower	152	4	39.25	pital	58	15	23.50
Beckenham Spire	169	16	1.25	East Cupola of ditto	58	45	37.12
Leith Tower (Centre)	169	39	15.00	Bermondsey Church	58	51	44.00
Croydon Church (Centre)	174	29	33.75	St. Saviour's Church (Centre)	59	30	25.87
Norwood Mill	8	27	53.12	Spire North end of London Bridge	61	38	45.87
Lewisham Church Vane	23	32	37.37	St. Paul's	62	5	30.87
Leigh Spire (Centre)	31	37	34.00	St. Dunstan's East	62	23	20.00
Peckham Chapel Spire	42	39	4.87	The Monument	62	6	22.50
Deptford Spire	<b>5</b> 0	21	9.56	The Tower Flag-staff	62	34	18.12
New Chapel east side of Black-				Flamsteed's Observatory, N.E.			
heath Cross	47	37	49.50	Cupola of the Tower	62	41	59.75
Centre of Octagon Room, Royal				St. Matthew's Bethnal Green, a			
Observatory	52	53	30.62	Tower with a Cupola	74	49	57.12
Vane Cupola of Greenwich Ob-				Limehouse Church	75	7	4.87
servatory	52	<b>5</b> 8	1.12	Charlton Church Flag-staff	81	14	6.37
Greenwich Spire	57	7	5.00	Direction of Roy's Station	98	15	0.00
Newington Butts Church Cupola	56	26	27.50	Woolwich Church Flag-staff	125	27	18.50
St. George's Church Spire (Cop-				,			
per Ball)	56	34	19.12				

## At Chingford.

	Vrotham and	Sever	ndroog.		S	everndroog a	nd St.	Paul's.	-
Observers.	Five Microscopes.	Mean. Five.	Two Microscopes.	Mean. Two.	Observers.	Five Microscopes.	Mean. Five.	Two Microscopes.	Mean. Two.
K. G. G.	16 34 61.20 63.55 16 34 61.80 60.60 16 34 60.10 61.85		59.50 62.00 63.87 65.75 60.38 61.25 64.87 62.37	61.99 60.75 64.81 60.81 63.62 62.40	K. C. & G. K. G. G.	39 0 37.05 36.20 36.33 34.75 38.00 37.60 39 0 34.70 35.10 34.50 36.90 Mean	36.62 35.54 37.80 34.90 35.70 36.11	35.13 35.13 34.94 34.00 23.13 37.00 34.02 33.25 36.25	35.13 34.47 30.06 34.31 34.75
W	Vrotham and	Severi	ndroog.		Severndi	roog, and Gre	enwic	h Observa	atory.
Observers.	Three Microscopes.	Mean. Three.			Observers.	Three Microscopes.	Mean. Three.		
К. С. К. К. К.	65.00	62.54 63.08 61.87 62.63 62.79 62.13 58.33			К.	12 51 23.92 27.34 24.50 24.66 26.00 23.83 26.17 24.58 12 51 27.59 23.67 Mean	25.63 24.58 24.91 25.37 25.63		
	Mean	61.91						,	

### Chingford (Continued).

With the Ordnance Theodolite, July 1323.

S	everndroog a	nd St.	Paul's.	Wroth	am, and Cen	tre of	Observat	ory.
Observers.	Three Microscopes.	Mean. Three.	-	Observers.	Three Microscopes.	Mean. Three.		
К.	39 ó 36.92 35.50	" 36.21		K.	29 26 25.96 29.09 27.33	27.52		
Westm	ninster Abbey	and S	Severndroog.	*	27.41 28.91 26.50	27.37 27.70		
Observers.	Microscopes.  42 52 09.42				28.67 27.75 28.75 24.92	28.21 26.82	1.50	
À	09.50 09.42 42 52 12.12				29 26 23.42 24.50	23.90		
	42 52 10.35 Mean 10.16				Mean	26.93		

### Observations for identifying the Station.

	Readings.		Readings.
St. Paul's	32 49 39.37	Hadleigh Church Beacon	108 33 59.37
Holloway Chapel	43 36 31.87	Barnet Church Vane	104 25 22.50
Chingford Church Tower	41 44 15.75	- W	

#### At Wrotham.

Cr	owborough, a	nd Le	ith Hill.			Fairlight and	Severi	ndroog.	
Observers.	Five Microscopes.	Mean. Five.	Two Microscopes.	Mean. Two.	Observers.	Five Microscopes.	Mean. Five.	Two Microscopes.	Mean. Two.
K.(A.M.)	49.70 49.95 53 57 52.45 53.40 53.70 51.25 51.25	51.4%	50.00 51.25 50.25 50.37 51.75 51.75 50.75 47.37 49.75 50.50 50.88 51.87 51.13	50.62 50.31 51.56 51.25 49.06 50.12 51.50	K. K. G. G. G.	47.40 45.65	46.52 47.07 45.72	45.87 50.75 46.75 46.38 48.75 46.88 46.63 43.75 49.25 47.12 48.75 48.50 44.87	48.31 46.56 47.81 45.19 48.18 48.62 46.43
G.	52.70		52.75		-	Mean	47.36		47.30
	Mean	51.13		50.63		-		1)	<del>'</del>

## Wrotham (Continued).

- , S	Severndroog a	nd Ch	ingford.	,	C	rowborough a	and Cl	hingford.	
Observers.	Five Microscopes.	Mean. Five.	Two Microscopes.	Mean. Two.	Observers.	Five Microscopes.	Mean. Five.	Two Microscopes.	Mean. Two.
G. G.	13 58 44.55 47.20	″ 45.87	42.38 47.37	" 44.87	G.	133 23 26.65 23.75	" 25.20	27.00 22.75	" 24.87
	13 58 41.00 45.30	43.15	38.62 39.88	39.25		Mean	25.20		24.87
C. G.	42.12 45.25 45.60	43.68	38.75 42.38 43.38	40.56	St	ede Hill and	Crow	borough.	1
G.	45.60 42.50	45.60 42.67	41.37 38.87	42.37 39.31	Observers.	Five Microscopes.	Mean. Five.	Two Microscopes.	Mean, Two.
с.	42.85 42.65		39.75 39.63	1					
G.	$\begin{array}{c} 42.50 \\ 13 \ 58 \ 46.40 \end{array}$	42.57 45.85	39.63 45.75	39.63 45.06	К. (р.м.)	51.30	51.47	50.88 51.50	51.19
c.	45.30 $44.85$ $44.40$	44.62	44.37 44.50 44.00	44.25	К. (р.м.)	51.95 49.45 49.70	49.57	51.82 50.87 48.75	49.81
G.	44.50 43.05	43.77	44.25 45.88	45.06	К. (а.м.)	47.65 47.55	47.60	47.06 47.88	47.47
-	Mean	44.20		42.26	К. (а.м.)	47.55 48.10	47.82	47.88 47.75	47.81
L	eith Hill and	Sever	ndroog.		C. (A.M.)	48.05	49.17	49.37 48.13	48.75
Observers.	Five	Mean.	Two Mi-	Mean.	C. (P.M.)	93 16 48.30 50.50 50.90	49.40	49.50 $51.88$ $51.25$	50.69
	Microscopes.	Five.	croscopes.	Two.	G. (P.M.)	49.20 52.85	50.05	51.00 53.75	51.12
К. (р.м.)	65 26 47.65 48.80	" 48.22	46.62 47.63	47.12	G. (P.M.)	48.60	50.72	51.25 51.88	52 50
K. (P.M.)	47.80 46.50	47.15	46.00 44.63	45.31	C. (P.M.)	52.00 50.55	51.82	53.63 51.63	52.75
С. (р.м.)	48.65 46.60	47.62	48.12 45.38	46.75	G. (P.M.)	51.90	50.70	56.00 52.00	motion 51.81
K. (A.M.) G. (P.M.)	49.25 46.40 65 26 48.55	47.82	47.87 45.88 47.25	46.87	G. (P.M.)	49.50 51.25 54.30	52.77	51.63 52.63 54.00	53.31
С. (р.м.)	44.65 44.80	46.60	45.50 47.13	46.37	к.	93 16 53.00 49.18	51.09	49.75	49,53
G. (P.M.)	49.10 47.10	46.95	48.50 47.75	47.81	к.	44.30 45.80	45.05	44.63 43.62	44.12
G. (P.M.)	49.50 52.00	48.30 48.79	48.12 51.00	47.93 48.93		46.35 $46.20$	46.27	44.50 45.50	45.00
К.	45.58 65 26 45.40	20.79	46.87 47.13	10.30		Mean	49.54		49.70
-	Mean	47.68		47.14	· · · · · · · · ·		······································	·	

#### Wrotham (Continued).

	Stede Hill an	d Fair	rlight.		F	airlight and (	Crowb	orough.	
Observers.	Five Microscopes.	Mean. Five.	Two Mi- croscopes.	Mean. Two.	Observers.	Five Microscopes.	Mean. Five.	Two Microscopes.	Mean. Two.
K. (P.M.) K. (P.M.) K. (A.M.) C. (A.M.) G. (P.M.) G. (P.M.) C. (P.M.)	50 5 42.35 38.55 39.85 40.50 37.67 37.75 39.05 39.95 38.95 50 5 41.90 42.90 42.90 43.10 41.75	42.42	42.13 38.00 40.12 38.25 36.44 37.25 37.12 38.62 38.00 41.87 45.00 40.88 43.13 42.50 41.37 41.75	"40.06 39.18 36.84 37.18 38.31 43.43 42.00 41.93		43 11 9.60 9.20 9.98 9.80 9.05 10.35 9.10 43 11 6.40 6.00 10.00 6.30 9.75 6.85 43 11 10.55	9.40 9.89 9.42 9.72 6.20 8.15 8.30 9.07	10.75 10.50 10.62 10.63 10.63 10.75 10.13 7.63 6.88 10.37 7.87 11.25 9.88 9.13 7.00 10.63	"10.62 10.62 10.63 10.44 7.25 9.12 10.56 8.06
G. (P.M.)	40.25	40.20	41.75 42.88 42.75 46.63	42.31 44.69	C (235)	supposed 11.85 motion. 14.40 11.35 7.80	0.57	10.63 12.87 8.88 7.25	11.25 8.06
-0.00	supposed 38.70 motion. 40.65	39.67	41.00 43.13	42.06	3	Mean	-	1.20	9.66
G. (р.м.)	40.55 41.70	41.12	43.12 44.38	43.75		Reject <sup>g</sup> 13.12	8.86		9.48
	Mean	40.62		40.98		-	<u>'</u>		

### Observations for identifying the Station.

	R	eadi	ings.		F	eadi?	ngs.
				A Tower about 2 miles distant			
A Tower near	23	51	18.25	Wadhurst Spire	88	12	1.00
The Station is in the north-	west	corn	er of a fic	eld upon Wrotham Hill, called "The Plain	s."		

#### At Leith Hill.

H	Ianger Hill a	nd St.	Paul's.		s	t. Paul's and	Sever	ndroog.	
Observers.	Five Microscopes.	Mean. Five.	Two Microscopes.	Mean. Two.	Observers.	Five Microscopes.	Mean. Five.	Two Microscopes.	Mean. Two.
G. K. G. K.	19 21 58.90 56.25	60.07 57.57	61.25 63.00 61.25 61.00 58.18 54.25 59.56 58.75	" 62.12 61.12 56.21 59.15	G. K. G. K.	16 1 17.00 14.70 15.80 15.05 16 1 14.60 17.65 13.32 17.90	15.42 16.12	17.62 14.13 15.50 14.12 15.13 19.38 14.50 19.00	15.87 14.81 17.25 16.75
	Mean	59.50		59.65		Mean	15.75		16.17

# Leith Hill (Continued).

Н	ınger Hill an	d Seve	erndroog.		V V	Vrotham and	Crowl	oorough.	->
Observers.	Five Microscopes.	Mean. Five.	Two Mi- croscopes.	Mean. Two.	Observers.	Five Microscopes.	Mean. Five.	Two Microscopes.	Mean. Two.
G.	35 23 14.80 12.35	" 13.57	16.50 14.87	15.68	G.	38 56 57.45 56.32	" 56.88	59.87 56.62	" 58.24
C.	12.75 14.85	13.80	15.00 15.38	15.19	G.	56.00 53.30	54.65	55.38 57.38	56.38
K. G.	11.75 35 23 12.60		13.75 12.88		G.	54.50 59.30	56.90	57.00 59.31	58.15
К,	11.90 13.88	12.25	13.00 14.69	12.94	G. G.	57.60 56.61	- 4-	57.00 56.36	
G.	12.95	13.41	12.75	13.72		54.70	55.65	56.25	56.30
	10.97 10.60	10.78	10.75	11.31	K.	55.65 55.55	55.60	59.00 55.62	57.31
K. G.	35 23 16.65	16.52	16.62 18.87	18.00	G.	55.85 53.90	54.87	55.00 56.75	55.87
<b>K</b> .	16.40 16.65	15.50	17.13 16.75	15.93	K.	38 56 56.75 57.05	56.90	56.75 57.37	57.06
G.	14.35 35 23 13.50	13.70	15.12 13.31		C.	56.85 56.85	56.85	56.13 58.25	57.19
K.	13.90 14.02		13.63 14.06	13.47	G.	56.90 55.55	56.22	56.87 54.75	55.81
	16.60	15.31	17.75	15.90	G.	38 56 55.00 55.00	55.00	55.25 53.88	54.56
	Mean	13.87	•••••	14.68		Mean	55.95		56.69
S	everndroog a	nd W	rotham.		Ш		)		1
Observers.	Five Microscopes.	Mean. Five.	Two Mi- croscopes.	Mean. Two.	118	nger Hill and Five	1	11	120
		- " ·			Observers.	Microscopes.	Mean. Five.	Two Mi- croscopes.	Mean. Two.
G.	28 7 17.57 19.05	18.31	18.56 17.87	" 18 <b>.</b> 21	G.	17 42 36.70	" 27 00	38.87	" 90.01
К.	28 7 17.60 17.63	17.61	18.75 17.63	18.19	G.	37.70 17 42 39.39	37.20	38.75 39.88	38.81
C.	17.35 16.40	16.87	17.50 17.00	17.25	G.	29.93 17 42 33.10	34.66	31.31 35.25	35.59
G.	16.35 15.05	15.70	17.38 16.19	16.78	К.	35.35 35.67	34.22	35.62	35.43
K.	28 7 17.75 13.40	15.57	16.25 11.63	13.94	G.	33.90 37.85	34.78	35.62 39.75	35.80
к.	1060	17.62	16.62	16.00	K.	40.75 36.45	39.30	40.63	40.19
G.	00 7 15 50	15.47	14.00	14.18	G.	37.75 17 42 37.65	37.10	37.50 39.12	38.31
к.	19 75	14.02	1412	13.75	к.	36.00	36.82	35.18 35.13	35.15
G.	00 7 10 15	16.57	17.05	16.06	Α.	37.97 39.75	38.86	37.31 38.25	37.78
		16.40	14.07			Mean	36.62		37.13
	Mean	16.42	• • • • • • • •	16.04			!		

### Leith Hill (Continued).

Observers.	Five Microscopes.	Mean. Five.	Two Microscopes.	Mean. Two.	Observations for identifying the Station.
G. G. K. G.	17 40 39.20 35.65 39.20 36.60 17 40 35.85 37.90	36.37 36.12 37.42 37.90 36.87 36.45	37.63 36.12 30.87 40.56 39.12 36.50 39.25 36.00 38.13 38.50 36.75 39.50	36.87 35.71 37.81 37.62 38.31 38.12	beyond Box Hill on high ground 59 6 50.1 Cupola Vane of Ewherst Church 90 17 1.1
	Mean	36.85		37.41	

The North-west Angle of Leith Tower is distant from the Station 66 feet 10 inches.



At Stede Hill.

F	airlight and	Crowb	orough.	* -	נ	Colsford and	Crowb	orough.	4
Observers.	Five Microscopes.	Mean. Five.	Two Microscopes.	Mean. Two.	Observers.	Five Microscopes.	Mean. Five.	Two Microscopes.	Mean. Two.
G.	53 13 24.15 24.25	" 24.20	25.15 26.38	" 25.76	G.	118 37 37.00 36.92	" 36.96	37.02 39.44	" 38.23
G. G.	53 13 23.35 25.30	24.33	23.62 25.25	24.44	G.	118 37 35.55 37.39		37.12 41.00 36.50	39.06
1	53 13 without clamping 53 13 25.65	****	22.38	-		118 37 without clamping 35.80	:	30.30	
G.	slight clamping25.00 23.95	24.47			G.	slight clamping 34.85 36.35	35.60		
	Mean	24.33		25.10	·	Mean	36.34	38.22	38.64

### Stede Hill (Continued).

	Fairlight and	Wro	tham.		/ (yee 1)	Tolsford and	d Fair	light.	
Observers.	Five Microscopes.	Mean. Five.	Two Microscopes.	Mean. Two.	Observers.	Five Microscopes.	Mean. Five.	Two Microscopes.	Mean Two
K. G.	97 58 15.72 17.02 97 58 15.60 17.80 97 58 14.38 out clamping. H Wrotham from 21.20 slight clamping 16.80 20.10 without clamping 18.20 15.85	Crowb 18.45		" 16.56 14.56 aching	G. G. K. G. K.	65 24 12.85 12.67 65 24 12.20 12.09 65 24 65 24 10.15 slight clamping 9.85 12.40 without clamping 11.15 12.10 slight clamping 65 24 11.15	" 12.76 12.14 11.12 11.62 10.72	11.87 13.06 13.50 16.12 14.12	" 12.46 14.81
	Mean	17.13		15.56	C.			11.32	
	Tolsford and	Wro	tham.			Mean	11.67	13.33	13.63
Observers.	Five Microscopes.	Mean. Five.	Two Mi- croscopes.	Mean. Two.	C	rowborough a	nd W	rotham.	
G.	163 22 28.57 29.67 163 22 27.80 29.89 163 22 29.85	29.12 28.84	27.43 30.62 27.12 31.25 28.50	29.02 29.18	Observers. G.	Five Microscopes.  44 44 51.57 52.75 44 44 52.25 52.50	Mean. Five. " 52.16	Two Microscopes.  50.41 51.18 50.00 50.25	Mean. Two. 50.79 50.12
G.	29.83 slight clamping	29.57 28.65		)	G.	44 44 slight clamping 51.80 56.15	53.97 52.83	52.00	50.45
1	Mean	29.52		29.10	i i	-			

# Observations for identifying the Station.

	$\mathbf{R}$	eadi	ngs.
Bottom of the Spindle of the Vane of Charing Church	$9\overset{\circ}{2}$	10	5.00
Bottom of the Spindle of the Vane of Lenham Church			
Bottom of the Spindle of the Vane of Harrietsham Church	128	24	39.37
Bottom of the Spindle of the Vane of Hollingbourn Church	51	15	6.12

## At Crowborough.

	Tolsford and	l Fairl	ight.		Leith Hill an	d Wro	otham.	
Observers.	Five Microscopes.	Mean. Five.		Observers.	Five Microscopes.	Mean. Five.	Two Microscopes.	Mean. Two.
G.	36 5 24 53 26.00	″ 25.26		K.	87 5 14.55 15.65	" 15.10	"	"
K.	28 48 25.60	27.04		K.	15 97 14.00	14.98	18.19 16.50	17.34
G.	36 5 20.75 23.40	22.07	*	G.	$14.83 \\ 16.05$	15.44	*	
G.	36 5 22.83 20.50	21.66		K.	15.40 14.30	14.85		
	Mean	24.01		C.	87 5 14.48 16.00	15.24		
,				G.	87 5 15.05 13.65	14.35	-	
40	Wrotham ar	id Tol	stord.	G.	87 5 13.90 12.30	13.10		
Observers.	Five Microscopes.	Mean. Five.		G. G.	87 5 87 5 16.90	1	14.38 12.87 17.25	15 05
G.	67 36 60.42	" 59.38		G.	14.60 87 5 17.90 15.20		17.25 17.63	17.25 16.63
к.	58.35 57.77 59.75	58.76	je.		87 5 16.15 13.60	1407	15.63 13.25 12.00	12.62
G.	60.67	60.43	:		Mean	·		15.96
<b>G.</b>	67 36 62.02 64.50	63.26		J	rant Church	and E	'airlight.	
	Mean	60.46				·	0	7
F	rant Church	and T	olsford.	Observers.	Five Microscopes.	Mean. Five.		
Observers.	Five	Mean.	,	K. K.	61 31 22.75 19.85	"		
	Microscopes.	Five.		G.	21.65 24.10	20.75		
G. G.	25 25 57.75 59.57	"	8 ,	K.	22.25	23.17		
K.	56.25	57.91			24.00 23.75	23.87	*	
	55.52 58.15	56.83	*	C.	61 31 22.15 24.75	23.45	 5	
G. G.	25 25 59.47 57.67 25 25 63.02	58.57		G.	$\begin{array}{cccc} 61 & 31 & 20.22 \\ & & 21.07 \\ 61 & 31 & 25.85 \end{array}$	20.64		
u.	Mean 58.42	57.77		G.	Mean	22.38		

## Crowborough (Continued).

	Leith 1	Hill an	d Sted	le Hill.			Wrotham ar	d Fair	rlight.	
Observers.		ive scopes.	Mean. . Five.	Two Microscopes.	Mean. Two.	Observers.	Five Microscopes.	Mean. Five.	Two Microscopes.	Mean Two
K. K. G. K. C. G.	129 4 129 4 129 4	34.40 36.85 35.00 35.91 36.35 35.75 35.55 36.78 39.10 33.30 35.58 32.70 35.42 37.25	35.92 36.13 35.65 37.94 34.44 34.06	38.62 38.25 33.38	38.43	K. G. K. C. G.	103 42 25.60 23.63 24.35 24.95 24.35 26.25 25.35 103 42 24.90 28.15 103 42 21.42 23.60 103 42 24.85 25.00 103 42 23.80	23.99 24.65 25.80 26.52 22.51 24.92 24.05	23.06 23.44 23.00	23.25
	Mea Wroth:	33.35 an		32.00 	32.69		Mean Stede Hill a	24.63	23.62 	23.28
Observers.		ive scopes.	Mean. Five.	Two Microscopes.	Mean. Two.	Observers.	Five Microscopes.	Mean. Five.	Two Microscopes.	Mean Two
K. K. G. K. C. G.	41 58	7 19.85 20.88 21.00 21.08 20.30 20.35 21.25 22.30 23.10 3 18.25 21.93 3 18.80	20.94 20.69 20.80 22.70 20.09 20.96	20.43 21.75	″ 21.09	K. K. G. K. C. G.	61 44 5.75 2.75 3.35 3.87 4.05 5.90 4.10 61 44 2.60 5.05 61 44 3.167 61 44 6.05 1.88	" 3.05 3.96 5.00 3.82 2.42 3.96	" 2.63 1.69	2.16
G.		23.12 21.10 19.75	20.42	20.13 20.00	20.06	G.	61 44 2.70 4.55	3.62	2.87 3.62	3.24

### Crowborough (Continued).

	rant Church a		JGC 22222		-	Stede Hill ar		
Observers.	Five Microscopes.	Mean. Five.	Two Microscopes.	Mean. Two.	Observers.	Five Microscopes.	Mean. Five.	
к.	0 12 43.00	"	"	"	G.	25 38 39.00	"	
К.	42.90 41.70	42.30	41.75 41.00	41.37	G.	38.03	38.81	
G. G.	41.25 39.77 41.80				K. G.	37.42 38.50 25 38 42.42	37.96	
к.	41.90 40.35				G.	38.27 25 38 43.22	40.34	
С.	0 12 40.45 40.30	10 27			<u></u>	41.38	42.30	
	Mean., 41.34	41.00		-		Mean	39.85	

### Observations for identifying the Station.

	Readings.		Readings.
Hartfield Spire	70 39 3.00	Crowborough Chapel Spire	11 31 19.00
Wadhurst Spire	7 4 48.75	Rotherfield Spire	23 52 44.25

### Observations repeated.

	Readi	ings.		R	eadi	ngs.
			Crowborough Chapel Spire			
Wadhurst Spire	120 53	48.00	Rotherfield Spire	137	41	46.50

#### At Tolsford.

	Blancnez and	d Fair	light.	Fien	nes, and Mor	ıtlamb	ert Statio	n.
Observers.	Five Microscopes.	Mean. Five.		Observers.	Five Microscopes.	Mean. Five.	Two Microscopes.	Mean. Two.
G.	118 34 46.95 48.65 118 34 48.55 118 34 51.72	" 47.80	_ =====================================	K.C.&G. K. K.C.&G.	17 30 15.40 14.70 15.80 17.13	" 15.05 16.46	17.46 16.41 16.38 17.83	16.93 17.10
	Mean 48.97	145		X	Mean	15.75	•••••	17.01

## Tolsford (Continued).

	Folkstone ar	nd Fair	light.	-		Fiennes and	l Fairl	ight.	
Observers.	Five Microscopes.	Mean. Five.	,		Observers.	Five Microscopes.	Mean. Five.	Two Mi- croscopes.	Mean. Two.
K.	136 51 43.40 46.05 45.20 46.83 136 51 46.85	46.01		15	K. C. & G.	113 18 17".08 16.72 17.50 18.53	16.90 18.01	1 <sup>"</sup> 9.82 18.87 20.34 22.17	" 19.34 21.25
K.	45.95 46.45	46.40 45.99				Mean	17.45	••••	20.29
G.	45.53 136 51 48.85 47.15	48.00			F	Cairlight and C	1	11	1
к.	50.45 46.65	48.55			Observers.	Five Microscopes.	Mean. Five.	Two Microscopes.	Mean. Two.
	Mean	46.61			G.	33 24 10.56 6.40	" 8.48	"	" "
-	Fairlight, and	1 14	Hill.	<u></u>	G. K.C.&G.	6.40 59.60 5.12	3.00	2.28	
Observers.	Five Microscopes.	Mean. Five.		3.	G.	33 24 10.00 8.38	9.19	9.62 8.31	8.96
G.	69 7 60.65 60.15	60.40				Mean 6.64	6.89	6.74	
К.	61.75 58.75	60.25			Cı	rowborough, a	and Ste	ede Hill.	
G. K.	69 7 57.50 60.60 59.33	59.05			Observers.	Five Microscopes.	Mean. Five.	,	
G.	59.70 69 7 54.50 54.95	59.51 54.72			G. G.	35 43 48.00 35 43 48.10	<i>"</i>		
к.	56.10 60.35	58.22			u.	55.35	51.72		
	Mean	58.69			-	Mean 50.48	<u> </u>	<u> </u>	
Folkstor	ne Church an	d Folk	stone Sta	tion.	<u>,                                    </u>	Tolkstone and	Notre	Dame.	T .
Observers.	Five Microscopes.	Mean. Five.	*		Observers.	Five Microscopes.			
G.	14 49 43.70 42.95	" 43.32		·	G. G. K.	10 44 15.90 10 44 16.20 17.20		· · · · · · · · · · · · · · · · · · ·	
	Mean	43.32			,	Mean 16.43			1

## Tolsford (Continued).

	Folkstone an	d Blar	ncnez.		5 (a > )	Nôtre Dame a	ınd Fa	irlight.	
Observers.	Five Microscopes.	Mean. Five.		o* 104.	Observers.	Five Microscopes.	Mean. Five.	Two Microscopes.	Mean. Two.
G.	18 16 56.45 57.40	" 56.92			G.	126 7 27.50 126 7 29.75	"	"	"
G.	18 16 57.13 Mean 56.99		-		K. K.C.&G.	30.30 32.68 33.88	33.28	35.00 37.59 38.42	38.00
	Nôtre Dame :	and F	iennes.			Mean 30.82		37.00	
Observers.	Five Microscopes.	Mean. Five.	Two Microscopes.	Mean. Two.	,	Stede Hill an	ıd Foll	kstone.	
к.	ı̂2 49 12.15	12.68	12.50	" 13.84	Observers.	Five Microscopes.	Mean. Five.		2
K.C.&G.	13.22 15.18 15.35	15.26	15.18 17.25 16.25	16.75		154 0 15.95 13.80	" 14.87		
	Mean	13.97	·	15.29	К. К.	13.92 15.16 13.05	14.54		
Nôtre	Dame, and I	Montla	ımbert St	aff.	G.	14.42 154 0 15.65 13.45	13.73 14.55		-
Observers.	Five Microscopes.	Mean. Five.	Two Mi- croscopes.	Mean. Two.	K. K.	13.70 14.22 14.77	14.49		
K.C.&G.	30 19 29.88 31.15	" 30.51	33.63 34.08	" 33.85	G. K.	154 0 16.65 17.90 13.45	17.27	*	
	Mean	30.51		33.85		13.00 Mean	13.22		
Fair	light, and Mo	ontlam	bert Staf	£ ;	Folk	stone, and To	1 (	en Churc	h
Observers.	Five Microscopes.	Mean. Five.	Two Microscopes.	Mean. Two.	Observers.	Five			
C.K.&G.	95 48 1.32 2.80 2.80 2.73	2.06 2.76	2.41 3.96 3.96 4.34	3.18 4.15	K.	Microscopes.  166 49 58.53 55.58 166 49 60.40			
	Mean	2.41		3.66		Mean 58.17		*	

### Tolsford (Continued).

Folksto	ne Pier Flag Stati		Folkstone	F	olkstone and	Crowb	orough.	
Observers.	Five Microscopes.	Mean. Five.			Five Microscopes.	Mean. Five.		
G.	13 4 14.75 13.40	14.07			170 15 55.25 46.75	″ 51.00		
<u> </u>	Mean				Mean	51.00		
Fair	light and Te	nterden C	hurch.	1	Nôtre Dame a	nd Bla	an <b>c</b> nez.	
Observers.	Five Microscopes.			Observers.	Five Microscopes.			
к.	29 58 8.75			G.	7 32 40.55			

## Angles derived at Tolsford.

Fiennes and Fairlight	$1\overset{\circ}{13}$	18	17.45	Folkstone and Fairlight	$1\overset{\circ}{3}6$	51	46.61
Fiennes and Montlambert	17	30	15.75	Blancnez and Fairlight (observed)	118	34	48.97
Montlambert and Fairlight.  Montlambert and Fairlight	95	48	1.70	Folkstone and Blancnez Folkstone and Blancnez (ob-	18	16	57.64
(observed)	95	48	2.41	served)	18	16	56.99
MeanMontlambert and Fair-				MeanFolkstone and Blanc-	<del>Plannin qua</del>		
light	95	48	2.05	nez	18	16	57.31
			***************************************		***************************************		
Folkstone and Fairlight	136	51	46.61	*			
Montlambert and Fairlight	95	48	2.05	Folkstone and Montlambert	41	3	45.35
Folkstone and Montlambert.	41	3	44.36	Folkstone and Blancnez	18	16	57.31
Folkstone and Nôtre Dame, Calais	10	44	16.43	Blancnez and Montlambert.	22	46	48.04
Nôtre Dame, Calais, and Fiennes.	12	49	13.97	Blancnez and Fairlight (observed)	118	34	48.97
Fiennes and Montlambert	17	30	15.75	Montlambert and Fairlight			
Folkstone and Montlambert.	41	3	46.15	Blancnez and Montlambert.	22	46	46.92
	41	3	44.56				48.04
MeanFolkstone and Mont-	,		Ser	MeanBlancnez and Mont-	P		***************************************
lambert	41	3	45.35	lambert	22	46	47.48
				•			

Tenterden Church and Folkstone Folkstone and Fairlight					0	1	11
Tenterden Church and Fair- light				the different parts of the circle from which they were obtained, that is as 30 to 10, we have,— MeanTenterden Church and			
light (observed)	<del></del>	58	8.75	Fairlight	29	58	10.86

### Observations for identifying the Station.

	R	eadi	ngs.		$\mathbf{R}$	eadi	ngs.
	0	1	11		0	1	11
Beachborough Summer-house	<b>5</b> 5	50	3.00	Right-hand Edge of the Summit			
Stanford Church	28	10	11.62	of Lyme Castle	166	56	25.37
Ashford Church	48	41	6.37	Left-hand Edge of the summit of			
Left-hand Edge of the Summit of				Saltwood Castle	107	14	4.50
Lyme Castle	166	49	5.00				

### At Folkstone Station.

		l	11	,	1		-	1	]
Observers.	Two Microscopes.		Two Mi- croscopes.		Observers.	Two Microscopes.		Two Mi- croscopes.	
A. K. C. G. A. K. A.	75 31 75 31 75 31 75 31 75 31 75 31 75 31 75 31 75 31		42.75 41.00 39.75 41.13 40.01 43.50 41.75 38.87		1821. A. K. K. K. A. K.	50 26 50 26 50 26 50 26 50 26 50 26 50 26		43.27 47.94 46.94 47.00 47.35 51.06	~
	Mean	••••	41.10		1	Nôtre Dame a	nd Bl	ancnez.	<u> </u>
Folksto	one Church V	ane,	and Tolsfo	ord.	Observers.	Five	Mean.	Two Mi-	Mean
Observers.	Two		Two Mi-			Microscopes.	Five.	croscopes.	Two
	Microscopes.		croscopes.		G.	9 21 19.85	" 18.50	18.75	17.50
К.	69 13		51.62		C.	17.15 9 21	18.30	16.25 21.12	17.50
	Mean		51.62			Mean	10.50	18.71	

### Folkstone Station (Continued).

	Fairlight and	l Tols	ford.		Blan	cnez and Mor	ıtlamb	ert Lamps.
Observers.	Five Microscopes.	Mean. Five.	Two Microscopes.	Mean. Two.	Observers.	Two Microscopes.	3	Two Mi- croscopes.
1821. C. 1822.	36 17 "	"	58.13	"	A. K. C.	25 1 25 1 25 1	-	44.75 48.31 47.81
G. G.	36 17 61.30 55.95	58.62	59.25 60.87 56.63	58.75	A. K. C.	25 1 25 1	motion	48.37 48.94
C. G. G.	36 17 36 17 36 17 56.35 54.65	55.50	54.75 54.56 55.13 54.25	54.69	G. A. C. K.	25 1 25 1 25 1 25 1		$\begin{array}{c c} 50.00 \\ 46.81 \\ 46.25 \\ 45.94 \end{array}$
	Mean	<b>57.</b> 06	56.70	56.72	C. A. K.	25 1 25 1 25 1		45.13 47.18 46.31
Dover	Station and N	lôtre I	Dame, Cal	lais.	A. K. A.	25 1 25 1 25 1		47.93 46.00 45.41
Observers.	Five Microscopes.	Mean. Five.	Two Mi- croscopes.	Mean. Two.		Mean	••••	46.69
G. G.	41 16 " 29.10 32.30	″ 30.70	25.81 29.88 32.13	″ 31.00	Dove	r Station and	Dove	r Flag-staff.
С.	41 16		32.00		Observers.	Two Microscopes.		Two Mi- croscopes.
	Mean	30.70	29.95		Α.	o í1		4.75
D	over Station	and B	lancnez.		K. K.	0 11 0 11		0.25 6.00
Observers.	Five Microscopes.	Mean. Five.	Two Microscopes.	Mean. Two.	A. K.	0 11 0 11		3.12 3.56
1821. A.	50 37 "	, <i>''</i>	48.12	"		Mean	••••	3.54
K. K.	50 37 50 37		47.19 53.00		Blancne	ez, and Dung	eness	Light-House.
A. C. A.	50 37 50 37 50 37	* 1	50.37 51.69 48.69		Observers.	Two Microscopes.		Two Microscopes.
K. K. 1822.	50 37 50 37		48.75 54.62		A. K.	100 33 100 33		27.50 29.31
G. C.	50 37 48.95 49.45 50 37	49.20	48.63 48.38 53.12	48.50	C. A. K.	100 33 100 33 100 33	į.	27.56 28.56 23.12
	Mean of all	49.20	50.23			Mean		27.21

### Folkstone Station (Continued).

	Blancnez and	l Fair	light.		Folkstone Pier Flag-staff and Tolsford.						
Observers.	Five Microscopes.	Mean. · Five.	Two Mi- croscopes.	Mean. Two.	Observers. Two Microscopes.			Two Microscopes.			
1821. A.	121 48 "	"	<b>1</b> 7.07	"	K.	84 37		<b>4</b> 9.0			
C.	121 48 121 48		08.37 15.06			Mean	••••	49.0			
K. 1822.	121 48		11.50		Dunge	eness Light-H	ouse a	nd Fairli	ght.		
G. C.	121 48 12.55 16.75 121 48	14.65	12.50 17.37 14.50	.37	Observers.	Two Microscopes.		Two Microscopes.			
		14.65	13.77		Α.	21 14		49.57			
	Mean used	14.13			C. A.	21 14 21 14	-	52.62 46.25			
Ŋ	Iontlambert a	nd Fa	irlight.			Mean	••••	49.48			
Observers.	Two Microscopes.		Two Mi- croscopes.		]	Nôtre Dame a	ınd Fa	airlight.			
A. C.	96 46 96 46		32.32 23.24		Observers.	Five Microscopes.	Mean. Five.	Two Microscopes.	Mean. Two.		
A. K. C. A.	96 46 96 46 96 46 96 46 96 46		27.88 25.50 28.82 25.12		G. G. C.	131 9 " 32.40 33.90	" 33.15	35.76 31.25 33.62 35.62	" 32.43		
C. A.	96 46 96 46		25.19 26.63 25.12			Mean	33.15	34.06			
Mean			26.65			2					
Rejectin	g the first and s	econd	26.32								

### Angles derived at Folkstone.

Montlambert and Fairlight				
Montlambert and Blancnez				
Blancnez and Fairlight	121	48	13.01	
Blancnez and Fairlight (observed)	121	48	14.13	
MeanBlancnez and Fairlight	121	48	13.57	

### Observations for identifying the Station.

	$\mathbf{R}$	eadi	ngs.		R	eadi	ngs.
	0	- 1	11		3	,	11
Dover Castle Flag-staff	175	51	39.50	Right-hand Summit of nearest			
Summit of the Left Side of Mar-				Martello Tower 9	9	38	36.25
tello Tower nearest the Beach	98	28	28.75	Top of the Flag-staff Folkstone			
Right-hand Summit of ditto	99	3	38.50	Harbour 11	9	46	45.25
Left-hand Summit of nearest Mar-				Folkstone Church Vane Spindle - 13	5	10	42.75
tello Tower	98	47	59.75	Tolsford Station 2	4	24	34.37

# At Fairlight.

	Stede Hill an	nd Tol	sford.		w	rotham and	<b>Folsfo</b>	rd Staffs.	
Observers.	Five Microscopes.	Mean. Five.	Two Microscopes.	Mean. Two.	Observers.	Five Microscopes.	Mean. Five.	Two Mi- croscopes.	Mean. Two.
1821. G. 1822. G.	45 27 45 27 45 27 45 27 54.95 57.00 55.92	55.97	53.56 55.56	"	G. G. G. C. G.	77 23 57.62 56.07 57.90 77 23 57.55 57.65 58.40 77 23 57.75	56.98 57.60 57.92	53.63	55.87
G. C. G.	45 27 56.17 53.62 51.45 51.87 45 27 54.65 57.50	54 89 51.66 56.07			W	Mean rotham and T	57.50	58.12 d Lamps	,
G.	45 27 50.20 53.75 55.85 55.70	51.97 55.77	51.12 55.81 57.25 56.87	53.46 57.06	Observers.	Five Microscopes.	Mean. Five.	Two Microscopes.	Mean. Two.
2	Mean	54.39	55.03	55.26	K. G.	77 23 58.50 57.35 61.65	57.92 60.26		
Observers.	Dungeness an	nd Bla	ncnez.		K. G.	58.88 77 23 60.25 56.18	58.21	F0.04	
к.	Microscopes.	77- 0-	croscopes.	7	K.	$\begin{array}{c} 77 \ 23 \ 58.35 \\ 60.35 \\ 58.10 \\ 61.00 \end{array}$	59.35 59.55	59.24 60.25 58.63 62.50	59.74 55.56
G. A. G. C.	4 59 4 59 4 59 4 59		12.12 15.13 12.68 13.41		C. <b>K</b> .	57.98 61.90 77 23 59.95 62.35	59.94 61.15	59.81 61.75 58.75 60.00	60.78 59.37
*	Mean		12.87	* 1		Mean	59.48	•••••	58.86

# Fairlight (Continued).

(	Crowborough	and St	ede Hill.			Crowborough	and F	olkstone.	
Observers	Five Microscopes.	Mean. Five.	Two Microscopes.		Observers.	Five Microscopes.	Mean Five.		Mear Two
G	65 2 35.86 32.87	" 34.36	,,,,	"	G.	117 20 51.25 47.07		"	"
G G	34.65	99 90	5		G.	48.62 47.57 117 20 46.85	48.09		,
C	37.98	37.27			C.	44.25 45.63	40.00		
G.	$\begin{array}{c ccccccccccccccccccccccccccccccccccc$	38.39	40.25 36.29	38.27		43.54 117 20	44.58	46.37	
G.	33.50 38.00	35.75	34.00 37.63	35.81		117 20 117 20 117 20		47.38 47.87 45.37	
-10.	Mean	35.83		37.04		117 20 117 20	- 27	43.88 47.75	
Folksto	one and Dung	eness l	Light-Ho	use.	G. G.	117 20 45.67 43.58	44.62	45.49 43.97	44.73
Observers.	Five Microscopes.		Two Mi- croscopes.		G.	44.80 46.90	45.85	45.37 46.75	46.06
G. K.	21 50		<b>6.51</b>			Mean	46.31	46.02	
C. G.	21 50 21 50		7.13 5.25		Bla	ncnez, and Fa	irligh	t Church.	
A. C. C.	21 50 21 50 21 50	**	12.75 4.50 2.25		Observers.	Two Microscopes.	95	Two Microscopes.	
A. K. G.	21 50 21 50 21 50		4.25 8.06 4.88		G. A.	6 16 6 16		28.00 27.81	
I			6.17		К. С.	6 16 6 16		29.63 26.63	
<del></del>	ng greatest and	least	5.80	[		Mean	••••	28.01	
Crowbor	ough Staff, a	nd Wr	otham La	amp.	Dungene	ss Light-Hou	se and	l Montlan	ıbert.
Observers.	Five Microscopes.	Mean. Five.			Observers.	Two Microscopes.		Two Microscopes.	
K. G.	33 6 32.18 31.65 29.60	31.91	y :	×	K. G. A.	22 38 22 38 22 38 22 38	-	42.13 35.25 40.50	
К.	22 6 21 05	29.34	ē.	. [	G. K.	22 38 22 38		38.87 36.50	
	Mean 30.42	30.62	-			Mean	••••	38.65	

## Fairlight (Continued).

	Tolsford and	Folks	tone.		Crowborough and Wrotham Staffs.					
Observers.	Five Microscopes.	Mean. Five.	Two Mi- croscopes.	Mean. Two.	Observers.	Five Microscopes.	Mean. Five.	Two Microscopes.	Mean. Two.	
· 1821. G. G. & C. K. & G. G.	6 50 " 6 50 6 50 6 50	"	1"4.69 11.38 14.94 16.75		G. G. G.	22 6 20 05	32.62 30.67	"	"	
1822. G. K.	6 50 20.44 17.20 18.58	18.82			C. G.	31.03 33 6 31.60 35.60	33.60	37.62 36.38	37.00	
G.	17.37 $18.05$	17.97	-			Mean 32.14	32.29			
G.	$\begin{array}{c} 18.92 \\ 6 50 17.25 \end{array}$	18.48 17.27			C	Crowborough a	and T	olsford.		
C.	17.30 $16.20$ $15.10$	15.65			Observers.	Five Microscopes.	Mean. Five.	Two Mi- croscopes.	Mean. Two.	
G.	$\begin{array}{c} 6 & 50 & 17.65 \\ & 19.80 \end{array}$	18.72			G.	110 30 30.81	<i>"</i>	"	"	
G.	6 50 14.57 13.95	14.26	14.12 11.87	12.99	G.	29.87 30.57	30.34 29.61	4		
G.	15.45 13.20	14.32	14.12 12.25	13.18	G.	28.65 110 30 29.60 26.95	28.27			
	Mean	16.94	13.77	13.08	C.	29.43 28.44	28.93			
Spire	for the Pole S	Star ar	d Folkst	one.		110 30 31.10 30.13	30.61	31.37 $27.90$	29.63	
Observers.	Five Microscopes.	Mean. Five.	Two Microscopes.	Mean. Two.	G.	29.35 33.70	31.52	31.25 34.50	32.87	
G.	62 50 55.15	"	"	"		Mean	29.88	•••••	31.25	
G.	62 50 54.10 53.75	53.92	54.74 53.62	54.18		Wrotham an	d Sted	e Hill.	-	
53.77	Mean 54.33				Observers.	Five Microscopes.	Mean. Five.	Two Mi- croscopes.	Mean Two	
Fair	light Church	and M	Iontlamb	ert.	G			"	"	
Observers.	Two Microscropes.		Two Microscopes		G G	. 31 56 1.38 4.03	2.71			
A. K.			58.19 61.50		G G		0 15	6.38 1.25	3.83	
	Mean		59.84			Mean 2.49	2.43			

## Fairlight (Continued).

	Wrotham an	d Tol	sford.		Tent	erden Church	and	Folkstone	٠.
Observers.	Five Microscopes.	Mean. Five.	Two Microscopes.	Mean. Two.	Observers. Two Microscopes. Two Microscopes.				
к. G.	77 23 58.50 57.35 61.65	″ 57.92		"	G. C. & A.	46 9 46 9		50.87 46.50	
G.	58.88 77 23 57.62	60.26			K. & G.	46 9		51.38-	
К.	60.25 56.18	58.21				Mean		49.58	
G. G.	56.07 57.90 77 23 57.55	56.98	*	- 17	Spire	for the Pole S	Star ar	nd Tolsfo	rd.
c.	57.65 58.40	57.60		*	Observers.	Five Microscopes.	Mean. Five.	Two Microscopes.	Mean. Two.
G. K.	77 23 58.35 60.35 58.10	59.35	59.24 60.25 58.63	59.74	G.	56 ó 37.90	"	"	"
C.	61.00 $57.98$	59.55 59.94	62.50 59.81	60.56 60.78	Ğ.	56 0 39.53 39.80	39.66	40.62 41.75	41.18
к.	61.90 77 23 59.95	61.45	61.75 58.75	59.38		Mean. 39.08			
G.	62.95 57.75 58.10	57.92	60.00 53.63 58.12	<b>55.88</b>	Frant and Stede Hill.				
	Mean	58.92		59.27	Observer.	Five Microscopes.		Two Mi- croscopes.	
E	Blancnez and	Montl	ambert.	-		51 18 "		19.18	
Observers.	Two Microscopes.		Two Microscopes.	-1	G.	51 18 51 18 13.53		16.13	
A. K.	17 39 17 39		26.00 31.13			Mean 13.53	••••	17.65	
G. A. G.	17 39 17 39 17 39 17 39		23.13 25.37 26.19		Fol	lkstone, and I	Fairlig	ht Churcl	ı.
	Mean		26.36		Observers.	Two Microscopes.		Two Mi- croscopes.	
Tenterden Church and Tolsford.		C. & A.	33 5		45.56				
Observer.	Two Microscopes.		Two Mi- croscopes.		C. A. K.	33 5 33 5 33 5		41.63 46.12 47.81	
G.	39 <sup>°</sup> 19 <sup>°</sup>		36.44			Mean		45.28	

### Fairlight (Continued).

Frant Church and Tenterden Church.				Wrotham and Folkstone.						
Observers.	Two Microscopes.	Two Mi- croscopes.	Observers.	Five Microscopes.	Mean. Five.	Two Microscopes.	Mean. Two.			
K. & G. C. A.	57 26 57 26 57 26	31.87 41.12 38.13	G. G.	84 14 14.82 14.12 16.82	" 15.47	# 2	"			
	Mean	37.04	C.	84 14 14.80 14.95 14.60	14.87					
	Frant and Crowbo	rough.	G.	84 14 13.20 11.30	12.25	7.75 10.37	9.06			
Observer.	Five Microscopes.			Mean 14.33	14.19					
• . G.	13 44 19.90									

## Angles derived at Fairlight.

Folkstone and Montlambert Folkstone and Tolsford				Folkstone Blancnez
Tolsford and Montlambert	51	19	1.68	Foll
Folkstone and Dungeness  Dungeness and Blancnez  Folkstone and Blancnez	4	59	12.87	Folkston Dungene
Folkstone and Blancnez, using the mean of all the angles between				Mean.
Folkstone and Dungeness Tolsford and Folkstone				Tenterde Tolsford
Tolsford and Blancnez	33	39	35.99	Tenter
				11

	0	,	"
Folkstone and Blancnez			
Blancnez and Montlambert	17	39	26.36
Folkstone and Montlambert.	44	28	45.03
Folkstone and Dungeness	21	50	5.80
Dungeness and Montlambert			
	44	28	44.45
	44	28	45.03
MeanFolkstone and Mont-			
lambert	44	28	44.74
Tenterden Church and Folkstone.	46	9	49.58
Tolsford and Folkstone			
Tenterden Church and Tolsford	39	19	32.64
Stede Hill and Frant Church	51	18	15.59
Crowborough and Stede Hill			
Frant Church and Crowborough	13	44	20.24

#### Observations for identifying the Station.

	Readings.				$\mathbf{R}$	eadi	ngs.
	0	,	"		0	,	"
Ashford Tower	176	. 0	7.62	East Edge of Fairlight Mill	130	<b>54</b>	24.62
Mr. Fuller's Observatory Dome	84	37	51.12	Hastings Church	4	9	9.12
Roy's Station	113	33	57	Dungeness Light-House	46	9	36.50
Church about three miles	131	32	13.25	Fairlight Church	57	25	13.49
West Edge of Fairlight Mill	149	55	59.25				

From the Station to the nearest angle of the Windmill, 69 feet 2 inches. Roy's Station from the nearest angle of the Windmill, 26 feet 4 inches. Roy's Station from the new Station, 87 feet  $8\frac{1}{4}$  inches.

#### At Blancnez.

Montlambert Lamp and Fairlight Lamp.	Montlambert Lamp and Folkstone Lamp.	Montlambert Staff and Fiennes.	Folkstone Lamp and Tolsford.
Two Microscopes.	Two Microscopes.	Two Microscopes.	Two Microscopes.
75 56 24.64 75 56 23.98 75 56 23.03 75 56 25.25 75 56 25.53	107 18 56.17 107 18 55.71 107 18 53.64 107 18 57.18 107 18 56.78	51 21 34.19 51 21 35.18 51 21 33.88 51 21 33.12 51 21 29.97 51 21 37.53	3 36 52.37 3 36 50.05 3 36 52.78 Mean 3 36 51.73
Mean 75 56 24.49	Mean 107 18 55.90	Mean 51 21 33.98	Folkstone Staff
Fairlight Lamp and Folkstone Lamp.  31 22 30.72 31 22 33.99 31 22 27.82 31 22 30.56 31 22 33.30 31 22 28.03 31 22 30.60 31 22 29.91 31 22 31.27 31 22 32.61 31 22 31.53 31 22 31.53 31 22 32.68 31 22 31.25	Dungeness, and Folkstone Lamp.  28 58 55.91 28 59 2.06 28 58 55.84  Mean 28 58 57.94  Dungeness, and Fairlight Lamp.  2 23 35.36 2 23 29.47 2 23 35.69	Dungeness, and Montlambert Lamp.  78 20 0.33  78 20 0.33  Nôtre Dame Calais, and Folkstone Lamp.  131 56 37.34 131 56 39.23 131 56 40.39 131 56 39.53	And Dover Station.  12 3 50.81 12 3 53.44 12 3 51.41  Mean 12 3 51.88 Add 2.00  12 3 53.88  Dover Station and Nôtre Dame Calais.  119 52 47.12 119 52 53.53
Mean 31 22 31.13	Mean 2 23 33.51	Mean 131 56 39.12	Mean 119 52 50.32

## Blancnez (Continued).

	7		Ţ
Folkstone Staff and Tolsford.	DoverCastleFlag-staff and Folkstone Staff.	Nôtre Dame Calais, and Fiennes.	North Foreland Light- House, and Folkstone Staff.
Two Microscopes.	Two Microscopes.	Two Microscopes.	Two Microscopes.
3 36 50.00 3 36 51.63 3 36 48.88 3 36 53.37	12 1 15.43 12 1 15.75 12 1 17.75 Mean 12 1 16.31	69 22 50.78 69 22 53.50 69 22 54.23 69 22 56.31	41 51 11.50 Add2.00 41 51 13.50
3 36 49.63 3 36 51.25 3 36 54.88 Mean 3 36 51.38	Add 2.00	69 22 52.92 69 22 52.79 69 22 53.63 Mean 69 22 53.45	Montlambert Lamp and Fiennes.
Subtract2.00  3 36 49.38	Dover Station, and DoverCastleFlag-staff.	Folkstone Staff, and South Foreland	51 21 30.67 51 21 27.04 51 21 29.85
Nôtre Dame Calais, and Folkstone Staff.	0 2 33.66 0 2 35.38 0 2 37.69	High Light. 17 54 17.50 Add 2.00	51 21 32.16 51 21 30.29 Mean 51 21 30.00
131 56 "44.94 131 56 39.22 131 56 37.93 131 56 37.43 Mean 131 56 39.88 Add 2.00 131 56 41.88	Mean 0 2 35.58  Tolsford, and Fairlight Staff.  27 45 41.86 27 45 42.62 27 45 37.13	17 54 19.50  Folkstone Lamp, and South Foreland High Light.  17 54 11.27	Centre of Dunkirk Tower, and Notre
Fairlight, and South Foreland High Light. 49 16 42.54	27 45 37.13 27 45 36.25 Mean 27 45 39.46	Folkstone Lamp, and South Foreland Low Light. 17 24 16.55	Mean 5 13 50.51  DoverCastleFlag-staff and Nôtre Dame Calais.
Folkstone Staff, and South Foreland Low Light.  17 24 16.37 Add 2.00	Tolsford, and Fairlight Lamp. 27 45 38.23 27 45 39.86 27 45 38.75	17 24 16.55  Fairlight, and South Foreland Low Light. 48 46 47.82	119 55 22.50 119 55 27.19 Mean 119 55 24.84
17 24 18.37	Mean 27 45 38.95	48 46 47.82	

#### Angles derived at Blancnez.

Montlambert and Folkstone	107	18	55.90
Folkstone and Tolsford $\frac{(3^{\circ} \ 36' \ 49''.38 \times 7) + (3^{\circ} \ 36' \ 51''.73 \times 3)}{10} =$	3	36	50.09
Montlambert and Tolsford	103	42	5.81
Fairlight and Folkstone			
Fairlight and Tolsford	27	45	41.04
Fairlight and Tolsford (observed)	27	45	38.95
MeanFairlight and Tolsford	27	45	39.99

The Station at Blancnez is in a right line drawn from the Ball of Nôtre Dame, Calais, to a point upon the wall of the Guard-house, distant from the north edge of the house 4 feet 3.25 inches, and from the north-east edge 14 feet 10.3 inches. Mr. Gardner cut a cross upon the wall, marking this point. The distance on the ground from a vertical line passing through the cross to the Station is 107 feet 5 inches.

#### At Montlambert.

	Folkstone Lamp and Blancnez.		Dungeness, and Folkstone Lamp.		Dungeness and Blancnez.
Observers.	Two Microscopes.	Observers.	Two Microscopes.	Observers.	Two Microscopes.
K. A. G.	47 39 18.25 47 39 19.44 47 39 21.00	A.	27 56 55.62 27 56 55.62	C. A. C.	
М.	47 39 19.75		21 30 33.02	Mean	75 51 14.25
C. A. K. A. A. A. & C.	47 39 20.75 47 39 18.87 47 39 16.93 47 39 17.19 47 39 17.25 47 39 15.50	Α.	Dungeness, and Fairlight Lamp. 10 47 54.25	С.	S. Foreland Light and Folkstone. 14 35 16.12
Mean	47 39 18.49		10 47 54.25		14 35 16.12
<b>A.</b>	Fairlight Lamp and Tolsford. 32 53 1.62	A.C.&K.	Fairlight Mill and Folkstone. 38 46 0.48	A. & C.	S. Foreland Light* and Folkstone.  14 58 42.5
	32 53 1.62		38 46 0.48	11. 0.	14 58 42.5

<sup>\*</sup> There are two Lights at the South Foreland, the High and the Low Light.

#### Montlambert (Continued).

	Fairlight Lamp and Folkstone Lamp.	* '	Tolsford and Blancnez.		Fiennes and Blancnez.
Observers.	Two Microscopes.	Observers.	Two Microscopes.	Observers.	Two Microscopes.
A. K. C. A. K. C.	38 44 55.94 38 44 57.44 38 44 50.38 38 44 52.44 38 44 54.50 38 44 52.82	C. C. A. & C.	53 31 8.79 53 31 12.00 53 31 6.44 53 31 9.08	K. G. M. C. K.	34 27 39.68 34 27 40.06 34 27 37.62 34 27 39.79 34 27 41.37
A. & C. C. A. K. C. A.	38 44 52.62 38 44 52.63 38 44 52.57 38 44 55.75 38 44 49.92 38 44 49.87 38 44 53.42	C. C. A. & C.	Fiennes and Tolsford. 87 58 48.58 87 58 50.50 87 58 47.36	K. A. C. K. C. C. A. A. & C. Mean.	34 27 39.75 34 27 41.07 34 27 39.18 34 27 40.31 34 27 38.68 34 27 38.50 34 27 40.88 34 27 40.92
C. A. & C.	Tolsford, and Folkstone Staff. 5 51 52.32 5 51 49.56	Mean	87 58 48.81		
Mean	5 51 50.94				

### Angles derived at Montlambert.

	0	~ 1	11
Fairlight and Folkstone	38	44	53.43
Tolsford and Folkstone	5	51	50.94
Fairlight and Tolsford	32	53	2.49
Fairlight and Tolsford (observed)	32	53	1.62
MeanFairlight and Tolsford	32	53	2.05
Fairlight and Folkstone	38	44	53.43
Folkstone and Blancnez	47	39	18.49
Fairlight and Blancnez	86	24	11.91

			11
Folkstone and Blancnez	47	<b>3</b> 9	18.49
Colsford and Folkstone	5	51	50.94
Tolsford and Blancnez	53	31	9.43
Tolsford and Blancnez(observed)	53	31	9.08
MeanTolsford and Blancnez	53	31	9.25

The Station at Montlambert is on the North Bastion, about  $7\frac{1}{2}$  feet from the angle, measuring from the foot of the parapet, and equally distant from the faces.

#### Observations of the Pole Star.

## 1821, October 3rd, at Blancnez. In the evening. Chronometer slow 1<sup>m</sup> 46<sup>s</sup>.4 on Mean Time.

	1	teadi	ngs.
Nôtre Dame, Calais	51°	49'	43".06
South Foreland High Light	117	47	13.55
South Foreland Low Light	117	17	18.63

Chronometer.	Mean Horary Angle.	Reading at the Star.	Mean Reading at the Star.	Azimuth.	Reading at the Meridian.
h m s 5 59 23.5 6 6 53.0	91 8 4	164 38 46.50 164 39 3.12	° ′ ″ 164 38 54.81	。 / // 2 35 59	° ′ ″ 162 2 55.81
6 13 10.0 6 16 43.0	88 10 34	164 39 10.25 164 39 12.00	164 39 11.12	2 36 14	162 2 57.12
6 18 35.0 6 20 5.0	87 3 46	164 39 11.37 164 39 11.62	164 39 11.49	2 36 13	162 2 58.49
6 21 21.0 6 23 16.0	86 19 46	164 39 9.87 164 39 8.75	164 39 9.31	2 36 10	162 2 59.31
		Telescop	e inverted.		
6 43 49.0 6 46 19.0	80 37 20	164 38 12.25 164 37 59.37	164 38 5.81	2 34 56	162 3 9.81

# 1822, August 25th, at Fairlight. In the morning. Chronometer slow 3<sup>m</sup> 7<sup>s</sup>.64 on Mean Time.

Chronometer.	Mean Horary Angle.	Reading at the Star.	Mean Reading at the Star.	Azimuth.	Reading at the Meridian.
h m s 8 58 43.50 9 1 11.50	94 19 42	135 47 3″.0 135 47 44.0	0 / " 135 47 37.5	0 / // 2 34 54	。 / // 138 22 31.50
		Telescop	e inverted.		
9 11 28.00 9 13 28.30	97 27 58	135 48 33.75 135 48 46.65	135 48 40.2	2 33 45	138 22 25.20
	Summit o	-		Reading. 19' 0".27	

# 1822, August 25th, at Fairlight. In the evening. Chronometer slow 3<sup>m</sup> 7<sup>s</sup>.64 on Mean Time.

Reading.

Wrotham Lamp...... 112° 55′ 34″.4

Chronometer.	Mean Horary Angle.	Reading at the Star.	Mean Reading at the Star.	Azimuth.	Reading at the Meridian
h m s 8 39 9 8 42 59 8 45 59 8 48 27	89 59 43.6 88 22 14	140 58 6.25 140 58 11.00 140 58 14.65 140 58 14.80	0 / " 140 58 8.62 140 58 14.72	0 / " 2 35 46 2 35 52	0 / " 138 22 22.62 138 22 22.72
	,	Telescope	inverted.		Y
8 52 20 8 54 25.7 8 58 15	86 49 28.4 85 36 16	140 58 18.60 140 58 16.65 140 58 8.55	140 58 17.62 140 58 8.55	2 35 49 2 35 44	138 22 28.62 138 22 24.55
0 00 10	3 9	3		eading.	138 22 24.33

# August 26th, at Fairlight. In the evening. Chronometer slow 3<sup>m</sup> 9<sup>s</sup>.9 on Mean Time.

Reading. Wrotham Lamp ...... 99° 36′ 33″.97

Chronometer.	Mean Horary Angle.	Reading at the Star.	Mean Reading at the Star.	Azimuth.	Reading Meridi	
h m s 8 39 42 8 41 51 8 44 4 8 46 22	89 2 12.50 87 55 24	127 39 18.20 127 39 21.35 127 39 24.05 127 39 22.40	127 39 19.80 127 39 23.22	0 / " 2 35 50 2 35 51	125 3	" 29.8 32.22
		Telescop	e inverted.			
8 49 50 8 51 54	86 30 25.50	127 39 18.05 127 39 14.45	127 39 16.25	2 35 49	125 3	27.25
8 53 55 8 55 22	85 33 38.50	127 39 11.65 127 39 8.80	127 39 10.22	2 35 44	125 3	26.22
	Wrotham	Lamp Much motio	99° in the Lamp.	Reading. 36' 35".58		* -

## August 27th, at Fairlight. In the evening. Chronometer slow 3<sup>m</sup> 12<sup>s</sup>.1 on Mean Time.

Reading.

Tolsford Lamp ...... 4° 48′ 41″.50

Chronometer.	Mean Horary Angle.	Reading at the Star.	Mean Reading at the Star.	Azimuth.	Reading at the Meridian.
h m s 8 36 59 8 39 20 8 40 49	88 39 23.50	135 27 38.00 135 27 28.80 135 27 36.95	° ′ ″ 135 27 33.40	。 / " 2 35 51	° ′ ″ 132 51 42.40
8 41 56 8 43 3	87 51 1.00 87 25 49.00	135 27 36.60 135 27 35.85	135 27 36.77 135 27 35.85	2 35 52 2 35 51	132 51 44.77 132 51 44.85
		Telescop	e inverted.		
8 45 19 8 49 23 8 50 44	86 21 09.00 85 30 15.00	$\begin{array}{ccccc} 135 & 27 & 14.80 \\ 135 & 27 & 28.45 \\ 135 & 27 & 22.55 \end{array}$	135 27 21.62 135 27 22.55	2 35 48 2 35 42	132 51 33.62 132 51 40.55
			100 27 22.00	Rea	ding. 35".08

## September 7th, at Tolsford. In the morning. Chronometer slow 5<sup>m</sup> 40<sup>s</sup> on Mean Time.

Reading.
Folkstone Staff ...... 34° 32′ 27″.43

Chronometer.	Mean Horary Angle.	Reading at the Star.	Mean Reading at the Star.	Azimuth.	Reading at the Meridian.		
h m s 7 25 27 7 28 32	% / " 84 26 42.50	116 29 33.45 116 29 27.40	。 , " 116 29 30.42	° ′ ″ 2 36 14	° ′ ″ 119 5 44.42		
7 30 42 7 36 46	85 53 4.50	116 29 22.10 116 29 15.45	116 29 18.77	2 36 25	119 5 43.77		
$7 \ 41 \ 14$ $7 \ 43 \ 48$	88 20 13.00	116 29 12.45 116 29 13.45	116 29 12.95	2 36 32	119 5 44.95		
	Telescope inverted.						
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	89 55 6	116 29 1.60 116 29 3.85	116 29 2.72	2 36 26	119 5 28.72		
7 51 24 7 53 25	90 49 0	116 29 7.55 116 29 11.95	116 29 59.7	2 36 20	119 5 29.75		
8 2 32 8 4 38.5	93 37 5	116 29 40.95 116 29 46.50	116 29 43.72	2 35 46	119 5 29.72		
·	Folkstone	Staff		Leading. 32' 16".25			

September 19th, at Crowborough. In the morning. Chronometer slow 2<sup>m</sup> 15<sup>s</sup> on Mean Time.

Chronometer.	Mean Horary Angle.	Reading at the Star.	Mean Reading at the Star.	Azimuth.	Reading at the Meridian.
h m s 6 59 30 7 1 24 7 2 37	88 44 47.00 89 17 23.00	129 54 14.55 129 54 17.60 129 54 19.05	0 / " 129 54 16.07 129 54 19.05	0 ' " 2 36 14 2 36 12	0 / " 132 30 30.07 132 30 31.05
		Telescop	e inverted.		
7 4 22 7 6 32	89 59 59.60	129 54 17.35 129 54 19.20	129 54 18.27	2 36 9	132 30 27.27
7 7 37 7 8 48	90 38 43.00	129 54 19.95 129 54 22.55	129 54 21.25	2 36 5	132 30 26.25
	Frant Church Staff				

# September 20th, at Crowborough. In the morning. Chronometer slow 2<sup>m</sup> 17<sup>s</sup>.6 on Mean Time.

Chronometer.	Mean Horary Angle.	Reading at the Star.	Mean Reading at the Star.	Azimuth.	Reading at the Meridian.
h m s 6 39 54 6 41 20	0 / " 84 45 38	129 54 46.40 129 54 44.55	0 / " 129 54 45.47	0 ' " 2 36 0	0 / " 132 30 45.47
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	85 23 14	129 54 41.60 129 54 38.40	129 54 40.00	2 36 5	132 30 45.00
6 45 37 6 46 40	86 8 43.50	129 54 35.50 129 54 34.15	129 54 34.82	2 36 10	132 30 44.82
6 47 55 6 49 9	86 44 42.40	129 54 32.60 129 54 30.90	129 54 31.75	2 36 13	132 30 44.75
		Telescop	e inverted.		
6 53 13 6 54 15	88 2 55.50	129 54 46.45 129 54 46.55	129 54 46.50	2 36 15	132 31 1.50
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	88 40 16.00	129 54 47.05 129 54 47.75	129 54 47.40	2 36 14	132 31 1.40
$\begin{array}{cccccccccccccccccccccccccccccccccccc$	88 56 16.00	129 54 49.10 129 54 50.45	129 54 49.77	2 36 13	132 31 2.77
7 1 46	90 3 45.00	129 54 54.55	129 54 54.55	2 36 9	132 31 3.55
	Frant Ch	urch Staff (steady)	120	Reading. 9 4' 34".01	

September 21st, at Crowborough. In the morning. Chronometer slow 2<sup>m</sup> 19<sup>s</sup>.6 on Mean Time.

	R	leadi	ng.
Frant Church Staff	 160°	14'	11".73

Chronometer.	Mean Horary Angle.	Reading at the Star.	Mean Reading at the Star.	Azimuth.	Reading at the Meridian.
h m s 6 41 32 6 44 11 6 47 11	0 ' " 86 36 33.60 87 41 37.00	98 4 25.55 98 4 22.30 98 4 20.90	98 4 23.92 98 4 20.90	2 36 12 2 36 15	0 / " 100 40 35.92 100 40 35.90
		Telescop	e inverted.		
$\begin{array}{ccc} 6 & 50 & 8 \\ 6 & 52 & 2 \end{array}$	88 40 16.40	98 4 31.45 98 4 32.20	98 4 31.82	2 36 14	100 40 45.82
6 53 42 6 54 47	89 27 46.40	98 4 33.80 98 4 34.15	98 4 33.97	2 36 11	100 40 44.97
	Frant Ch	urch Staff		eading. 14' 18".66	

### Deductions from the preceding Tables.

#### Blancnez.

1821, October 3rd:	Mean Reading at the M	[eridian				•	$1\ddot{6}2$	2	5 <sup>"</sup> 7.68
	At Nôtre Dame,								
Between Nôtre	Dame Calais, and the Mo	eridian					69	46	45.38
Between Nôtre	Dame Calais, and Fairlig	ght .	•	•	•	•	163	19	11.63
Between Fairlig	ght and the Meridian .		•	•		i ·	93	32	26.25
Between Fairlig	tht and the Meridian, using	ng the S	out	h I	or	e-			
land High Li	ght		•				93	32	26.67
Between Fairlig	tht and the Meridian, using	ng the S	out	h I	For	e-			
land Low Lig	ght		•	•	•	•	93	<b>32</b>	26.87
		N	<b>I</b> ea	n	•	•	93	32	26.60

8 ,	162 117		9.81 16.21
Between the South Foreland High Light and the Meridian Between the South Foreland High Light and Fairlight .			53.60 42.54
Between Fairlight and the Meridian (Telescope inverted) Between Fairlight and the Meridian, using South Foreland Low Light (Telescope inverted)			36.14 35.10
	93	32	35.62 26.60
Between Fairlight and the Meridian	93	32	31.11
Fairlight.			
1822, 25th August A.M. Reading at the Meridian			31.50 59.30
Between the Spire and the Meridian	$\frac{4}{62}$		32.20 54.33
Between Folkstone and the Meridian			22.13 18.67
Between Blancnez and the Meridian	85	36	40.80
	138 134		2 <sup>"</sup> 5.20 0.27
Between the Spire and the Meridian	$\frac{4}{62}$		24.93 54.33
Between Folkstone and the Meridian (Telescope inverted) Between Folkstone and Blancnez		•	29.40 18.67
	-	,	
Between Blancnez and the Meridian (Telescope inverted) Between Blancnez and the Meridian (above)			48.07 40.80
Between Blancnez and the Meridian (Mean)	85	36	44.43

LONGITUDE BETWEEN PARIS AND GREENWICH.		233
25th August P.M. Mean Reading at the Meridian Wrotham Lamp		
Between Wrotham and the Meridian		
Between Blancnez and the Meridian	85 36	44.65
Mean Reading at the Meridian (Telescope inverted)	$1\overset{\circ}{3}8 \ \overset{'}{2}2$ $112 \ 55$	26.58 33.00
Between Wrotham and the Meridian		
Between Blancnez and the Meridian (Telescope inverted) Between Blancnez and the Meridian (above)		
Between Blancnez and the Meridian Mean	85 36	41.97
26th August P.M. Mean reading at the Meridian Wrotham Lamp		
Between Wrotham and the Meridian		
Between Blancnez and the Meridian	85 36	35.83
Mean Reading at the Meridian (Telescope inverted)		
Between Wrotham and the Meridian		
Between Blancnez and the Meridian (Telescope inverted) Between Blancnez and the Meridian (above)		
Between Blancnez and the Meridian Mean	85 36	38.77

27th August P.M. Mean Reading at the Meridian Tolsford Lamp	
	51 56 57.50 33 39 35.99
Between Blancnez and the Meridian	85 36 33.49
Mean Reading at the Meridian (Telescope inverted)	132 51 37.08 4 48 35.08
	51 56 58.00 33 39 35.99
Between Blancnez and the Meridian (Telescope inverted) Between Blancnez and the Meridian (above)	
Between Blancnez and the Meridian Mean	85 36 33 74
Summary.	
At Fairlight, the Angle between the Meridian and Blancne	ez:
8 ,	85 36 44.43 85 36 41.97
August 26, P.M August 27, P.M	85 36 38.77 85 36 33.74
Between the Meridian and Blancnez	85 36 39.73 33 39 35.99
Between the Meridian and Tolsford	51 57 3.74 110 30 29.88
Between the Meridian and Crowborough	58 33 26.14

### Tolsford.

Toisiord.
1822. September 7th, A.M. Reading at the Meridian 119 5 44.38  Folkstone 34 32 27.40
Between Folkstone and the Meridian
Between Folkstone and Crowborough 189 44 6.50
Between Crowborough and the Meridian
Mean reading at the Meridian (Telescope inverted)
Folkstone
Between Folkstone and the Meridian 95 26 46.86
Between Folkstone and Crowborough 189 44 6.50
Between Crowborough and the Meridian (Telescope inverted) 94 17 19.64
Between Crowborough and the Meridian (above) 94 17 23.48
Between Crowborough and the Meridian Mean . 94 17 21.56
Crowborough.
September 19th, A.M. Mean reading at the Meridian 132 30 30.56
Frant Church 12 4 8.55
Between Frant and the Meridian 59 33 37.99
Between Tolsford and Frant 25 25 58.37
Between Tolsford and the Meridian
Mean reading at the Meridian (Telescope inverted) 132 30 26.76
Frant Church
Between Frant and the Meridian
Between Tolsford and Frant
Between Tolsford and the Meridian (Telescope inverted) . 84 59 37.78
Between Tolsford and the Meridian (above) 84 59 36.36
Between Tolsford and the Meridian Mean 84 59 37.07

September 20th, A.M.	Mean reading a	t the M	[eridi	an				$13\overset{\circ}{2}$	20	45.01
•	S	Frant								22.70
Between Frant an	d the Meridian							50	33	37 60
Between Tolsford		• •								58.37
	•			•						
Between Tolsford	and the Meridian	• •	• •	•	•	•	•	84	59	36.06
N/ 1' 41 - N/	(Touristic of (TD-1)	•	4 1					100	. !	"
Mean reading at the M	teridian (Telescop									2.30 34.01
									4	34.01
Between Frant an										31.71
Between Tolsford	and Frant	• •	• •	•	•	•	٠	25	25	58.37
Between Tolsford	and the Meridian	(Teles	cope :	inve	ert	ed)		84	<b>59</b>	30.08
Between Tolsford	and the Meridian	(above	e) ·		•	•		84	<b>59</b>	36.06
Between Tolsford	and the Meridian		Me	an	•		٠	84	59	33.07
September 21st., A.M.	Mean reading a	t the M	<b>I</b> eridi	an		•		$10\mathring{0}$	40	3 <sup>"</sup> 35.91
•	,									11.73
Between Frant an	d the Meridian						•	50	33	35.89
Between Tolsford										58.37
Between Tolsford	and the Meridian	• •	• •	•	•	•	•	84	59	34.19
Mean reading at the M	Ioridian (Taloscon		(bot					100	46	4" 00
		·	•							18.66
Between Frant an										•
Between Tolsford	and Frant	• •	• •	•	•	•	•	25	<b>2</b> 5	58.37
Between Tolsford		`	-			,				
Between Tolsford	and the Meridian	(above	e) .	•		•	•	84	<b>59</b>	34.19
Between Tolsford	and the Meridian		Mea	an			×I	84	59	32.91

### Summary.

At Crowborough the Angle between							
Tolsford—19th September, A.M					$8\overset{\circ}{4}$	59	37.07
20th September, A.M							
21st September, A.M	•			•	84	59	32.91
Between the Meridian and Tolsford .				•	84	59	34.35
Between Crowborough and Fairlight .			•		36	5	24.01
Between the Meridian and Fairlight .	•	•	•	•	121	4	58.36

